

Gravitational Waves Detection

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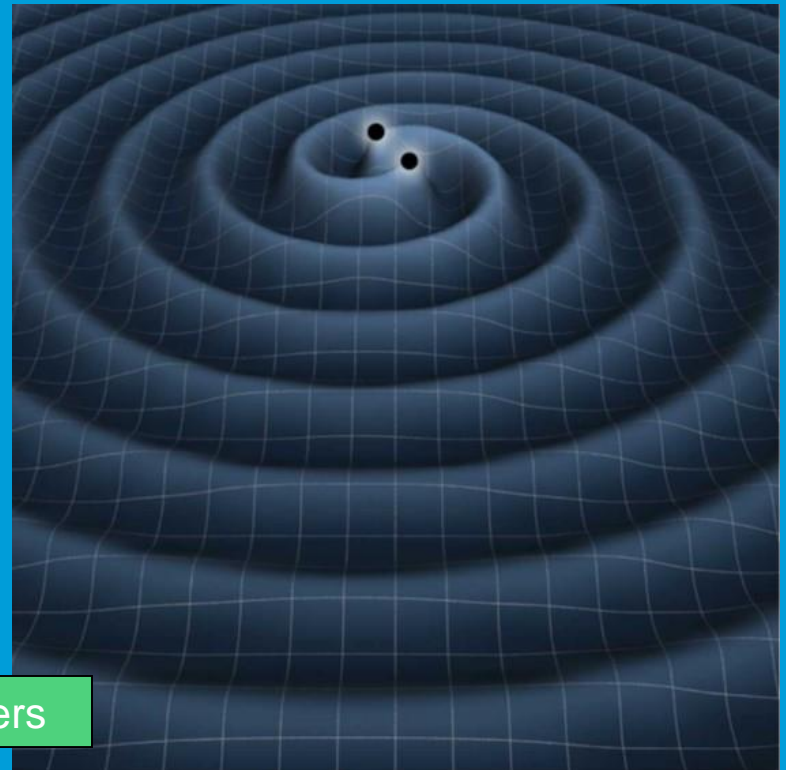
<http://pcgiammarchi.mi.infn.it/giammarchi/>



- Gravity as a fundamental force
- General Relativity
- Radiation of Gravitational Waves
- Indirect evidence (1974)
- Direct Detection (2016)

GW150914, GW151226, (LVT151012)

Disclaimer: the speaker is not one of the discoverers





Observation of Gravitational Waves from a Binary Black Hole Merger

B. P. Abbott *et al.**

(LIGO Scientific Collaboration and Virgo Collaboration)

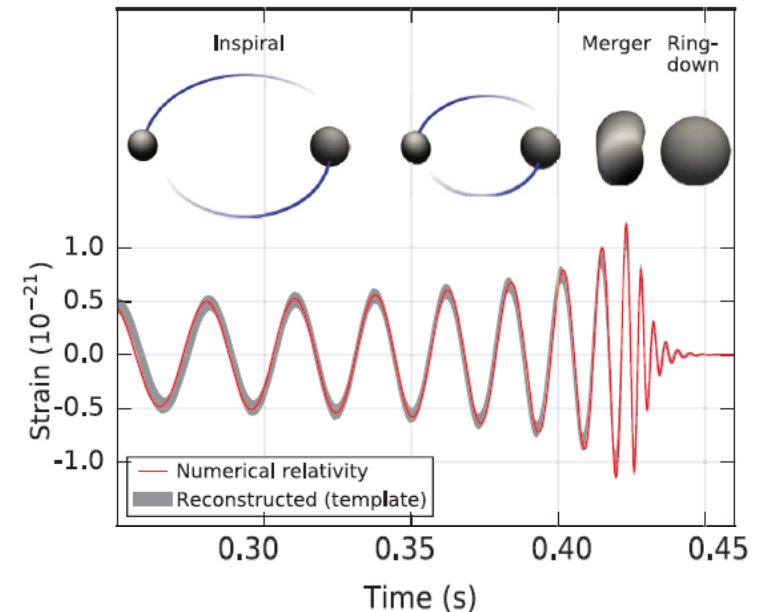
(Received 21 January 2016; published 11 February 2016)



11 February 2016 Press Conference
National Science Foundation
Washington DC - USA

VIII. CONCLUSION

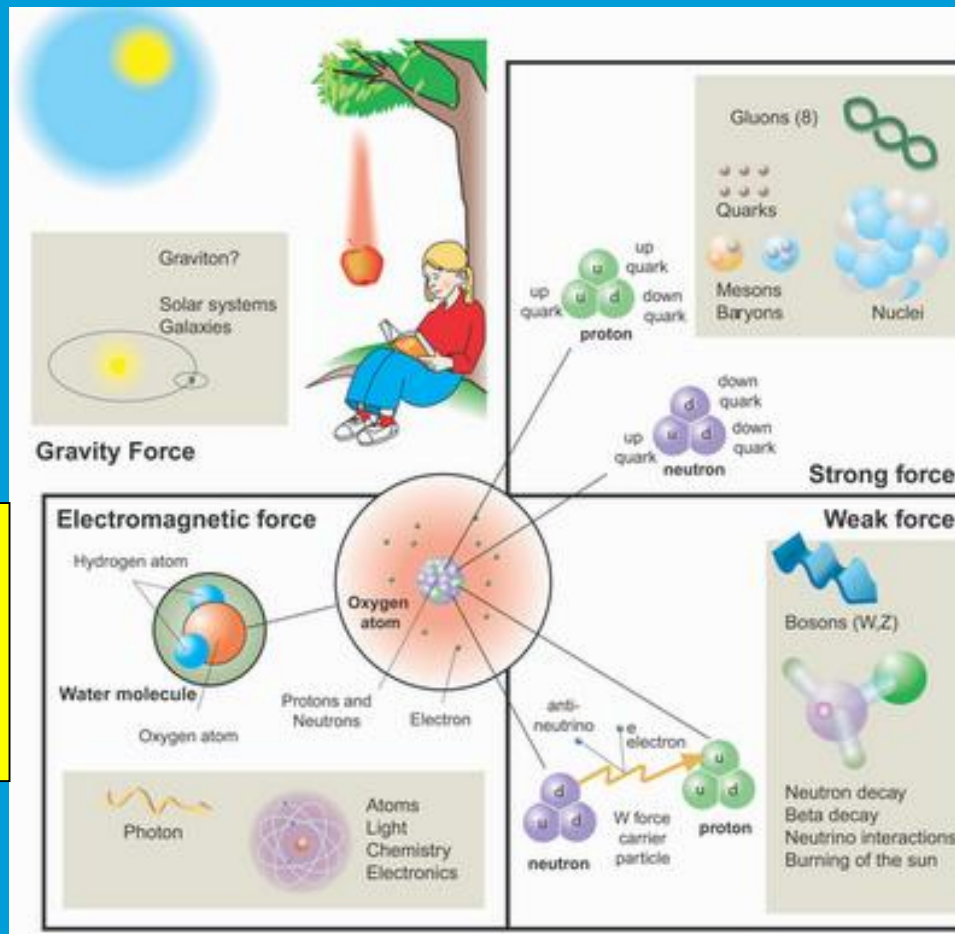
The LIGO detectors have observed gravitational waves from the merger of two stellar-mass black holes. The detected waveform matches the predictions of general relativity for the inspiral and merger of a pair of black holes and the ringdown of the resulting single black hole. These observations demonstrate the existence of binary stellar-mass black hole systems. This is the first direct detection of gravitational waves and the first observation of a binary black hole merger.



1. Gravity as a fundamental force

Gravitational Interaction: classical theory (A. Einstein) in 1915. Responsible for the stability of matter on macroscopic scales.

- Gravity as a fundamental force
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Strong Nuclear Interaction: a short range force, confined to a length of 10^{-15} m.

Elettromagnetic Interaction: takes place between charged particles (responsible for atomic stability).

Weak Nuclear Interaction: a subnuclear short-range (10^{-16} m) force.













Quark and Leptons: the fundamental matter constituents

Elementary (structureless) down to 10^{-18} m or more

Well defined charge and spin

Matter in ordinary (Earth-like) energy conditions

Ordinary Matter

| | LEPTONS | | QUARKS | |
|--|--|--|--|---|
| FIRST FAMILY "Ordinary" matter, least massive |  electron |  electron neutrino |  up |  down |
| SECOND FAMILY Similar properties, more massive |  muon |  muon neutrino |  charm |  strange |
| THIRD FAMILY Rarest particles, most massive |  tau |  tau neutrino |  top |  bottom |



M
a
s
s

Make up unstable particles

Decay to stable particles

These are the constituents. How do they interact between them?

The concept of elementary interaction

Newton

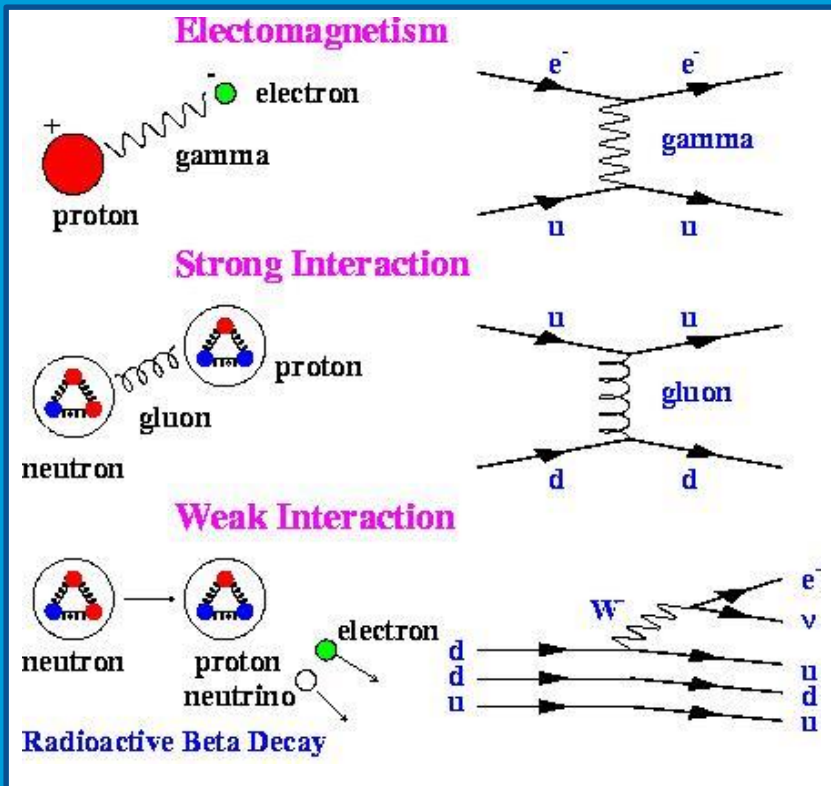
Action at a distance

Faraday, Maxwell

The classical field

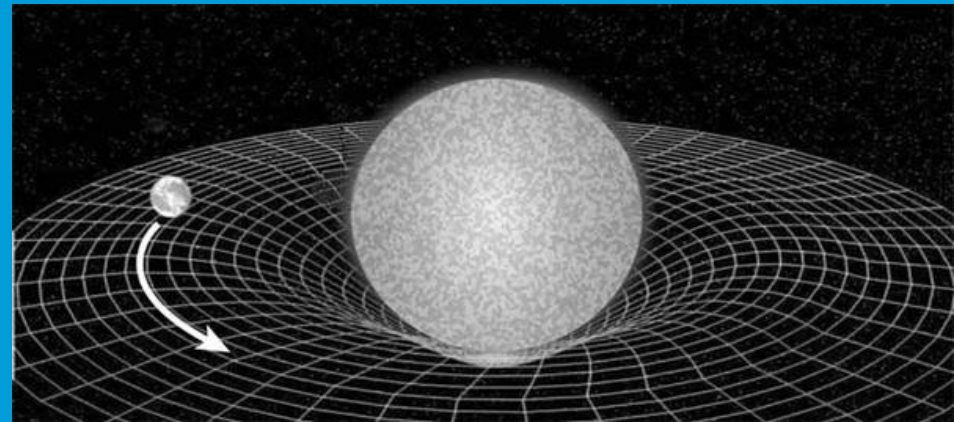
$$F = \frac{k}{r^2}$$

Quantum Fields (quanta exchange)



The only non-quantum force

Gravity (spacetime curvature)



Gravity is a special case :



- Gravity is way less intense than other forces on microscopic scales
- Gravity behaves the same for all bodies
- There is no Quantum Theory of Gravity
- Gravity is the decisive dynamical force in the Universe

$$G = \frac{2.12 \times 10^{15}}{kg^2} \hbar c$$

$$\alpha = \frac{e^2}{\hbar c} = \frac{(4.8 \times 10^{-10})^2}{1.054 \times 10^{-27} \times 3 \times 10^{10}} \frac{dyne cm cm}{erg s \frac{cm}{s}} = \frac{1}{137}$$

Comparing gravity and electromagnetism:

$$\frac{2.12 \times 10^{15}}{kg^2} \frac{mm}{r^2} \hbar c \quad ?? \quad \frac{\alpha}{r^2} \hbar c$$

$$\frac{2.12 \times 10^{15}}{kg^2} mm \quad ?? \quad \alpha$$

Need to choose charges and masses. For the case of two protons

$$\frac{2.12 \times 10^{15}}{kg^2} (1.67 \times 10^{-27} kg)^2 \quad ?? \quad \alpha$$

$$5.9 \times 10^{-39} \ll \frac{1}{137}$$

...gravity weaker by many orders of magnitude

2. General Relativity

Gravity concerns all forms of energy of the Universe (mass included)
Responsible of bounds between macroscopic bodies

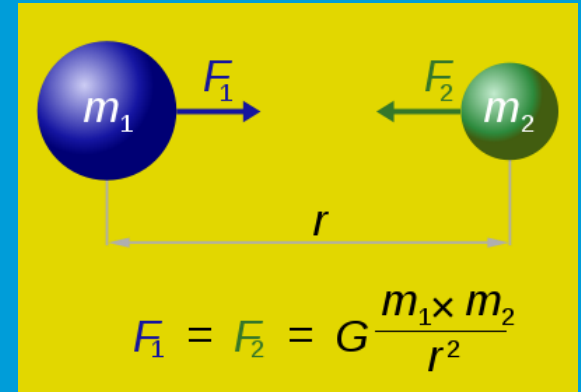
- Gravity as a fundamental force
- **General Relativity**
- Radiation of Gravitational Waves
- Indirect evidence (1974)
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Classical field theory (Newton, 1687) for the case of masses

$$\nabla^2 \phi = 4\pi G \rho$$

Gravitational potential

Mass density



“Geometrized” spacetime field theory (Einstein, 1915)

General Relativity

Principle of Equivalence between inertial mass (inertia to force) and gravitational mass (gravitational charge) → gravity as a property of the spacetime background

Far away from sources of mass/energy (in a flat spacetime)

$$g_{\alpha\beta}(x) \rightarrow \eta_{\mu\nu} = \begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{bmatrix}$$

Einstein Tensor

Energy-Momentum Tensor

Metric Tensor

$$G_{mn} = \frac{8\rho G}{c^4} T_{mn}$$

$$G_{\mu\nu} = G_{\mu\nu}(\partial_\theta \partial_\varepsilon g_{\alpha\beta})$$

Einstein's Field Equations

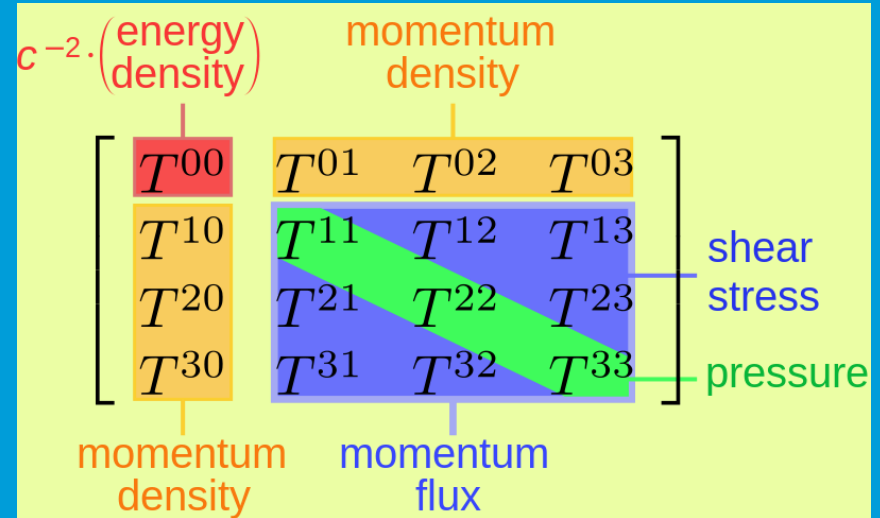
$$G^{\alpha\beta} = 8\pi T^{\alpha\beta}$$

Ten coupled differential equations

$$G^{\alpha\beta} = R^{\alpha\beta} - \frac{1}{2} g^{\alpha\beta} R$$

$$R_{\alpha\beta} = R^{\mu}_{\alpha\mu\beta}$$

$$R^{\lambda}_{\beta\mu\nu} = \frac{1}{2} g^{\lambda\sigma} (g_{\sigma\nu,\beta\mu} - g_{\sigma\mu,\beta\nu} + g_{\beta\mu,\sigma\nu} - g_{\beta\nu,\sigma\mu})$$



Determine the evolution of g based on initial conditions. They are not independent equations because of the Bianchi Identity

$$G^{\alpha\beta}_{;\beta} = 0$$

Six really independent equations to determine the six independent functions (among the ten g components) that characterize the geometry independent of coordinates

3. Radiation of Gravitational Waves

- Gravity as a fundamental force
- General Relativity
- **Radiation of Gravitational Waves**
- Indirect evidence (1974)
- Direct Detection (2016)

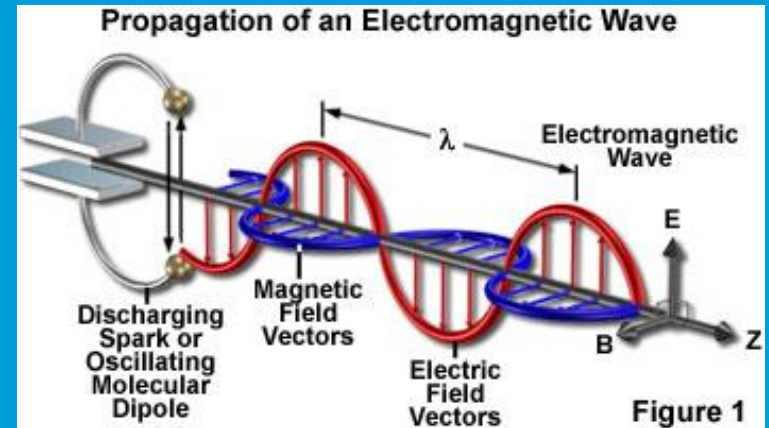
Electromagnetic radiator

- Oscillating Dipole
- E.M. waves discovered in 1886 (Hertz).

$$(\nabla^2 - \partial_t^2)A^\mu = 0$$

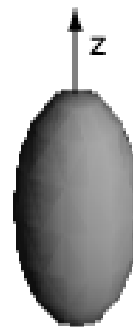
4-vector

- Two polarization states
- Photon: spin 1

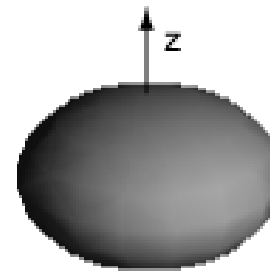


Gravitational radiator

Dipole gravitational radiation in GR is not present because of conservation of the linear momentum of N-body system



$Q > 0$
Prolate



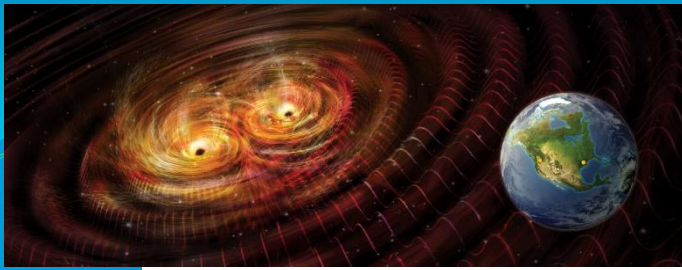
$Q < 0$
Oblate

Classical definition

$$Q_0 = \int \rho(3z^2 - r^2)dV$$

$$Q = \frac{3K^2 - I(I+1)}{(I+1)(2I+3)} Q_0$$

- Oscillating Quadrupole



- A Strong faraway Source (not treated in detail)
- A little perturbation (at the Earth)

GWs in linear gravity

- We consider **weak gravitational fields**:

$$g_{\mu\nu} \approx \eta_{\mu\nu} + h_{\mu\nu} + \mathcal{O}(h_{\mu\nu}^2)$$

↑
flat Minkowski metric

- The GR field equations in vacuum reduce to the standard **wave equation**:

$$\left(\frac{\partial^2}{\partial t^2} - \nabla^2 \right) h^{\mu\nu} = \square h^{\mu\nu} = 0$$

- Comment: GR gravity like electromagnetism has a “**gauge**” freedom associated with the choice of coordinate system. The above equation applies in the so-called “**transverse-traceless (TT)**” gauge where

$$h_{0\mu} = 0, \quad h^\mu{}_\mu = 0$$

Wave solutions

- Solving the previous wave equation in weak gravity is easy. The solutions represent “plane waves”:

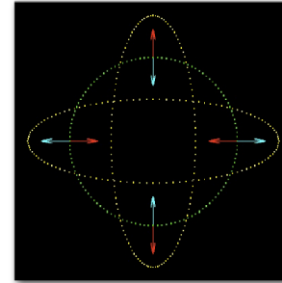
$$h_{\mu\nu} = A_{\mu\nu} e^{ik_a x^a}$$

↑
wave-vector

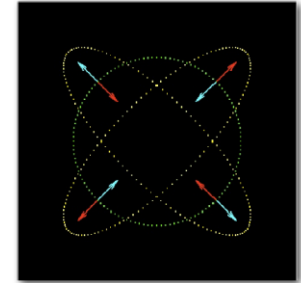
- Basic properties: $A_{\mu\nu} k^\mu = 0$, $k_a k^a = 0$
- ↑
transverse waves
↑
null vector = propagation along light rays

- Amplitude: $A^{\mu\nu} = h_+ e_+^{\mu\nu} + h_\times e_\times^{\mu\nu}$
- ↑
two polarizations

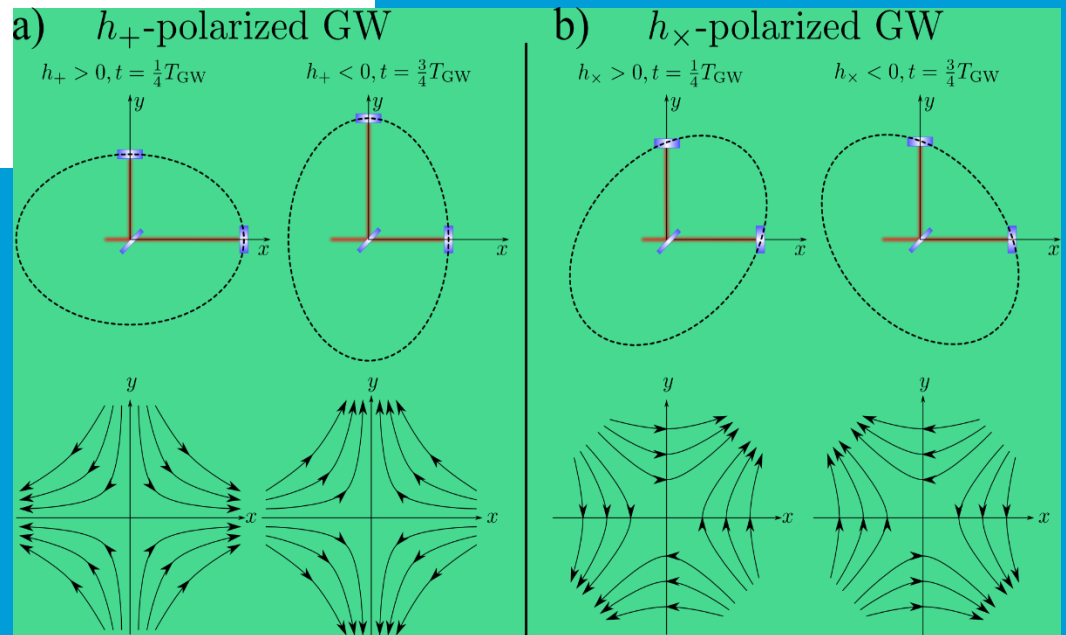
- GWs come in two polarizations:



“+” polarization



“x” polarization

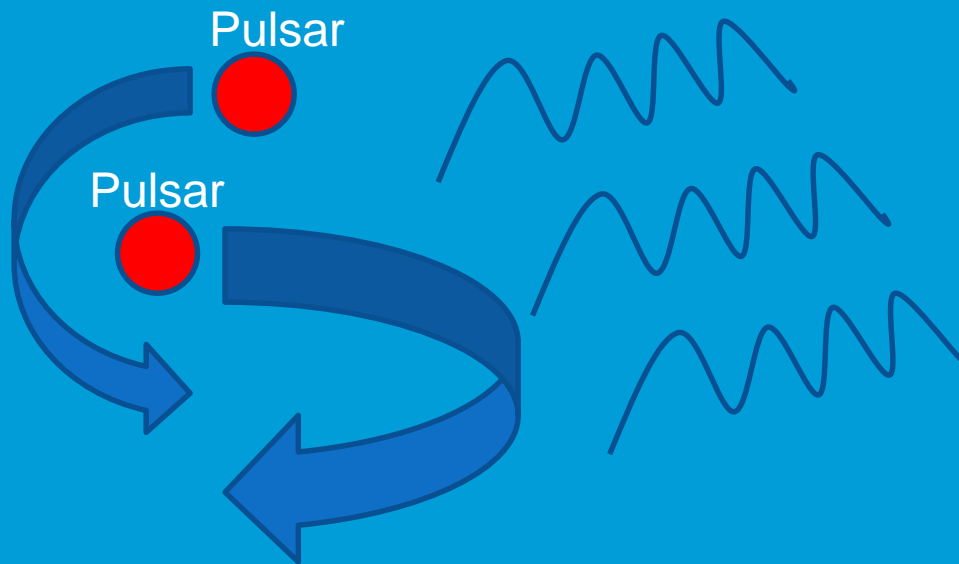


A Gravitational Wave entering the slide from above will alternatively deform the elliptical shape (where mirrors can be located)

4. Indirect evidence

Two massive and compact objects orbiting one around the other

- Gravity as a fundamental force
- General Relativity
- Radiation of Gravitational Waves
- **Indirect evidence (1974)**
- Direct Detection (2016)



The system can emit gravitational waves (if its quadrupole moment changes)

Energy conservation tells that the energy of the system decreases!
The orbit parameters change

The effect can be detectable!

INDIRECT evidence of the existence of Gravitational Radiation from the source

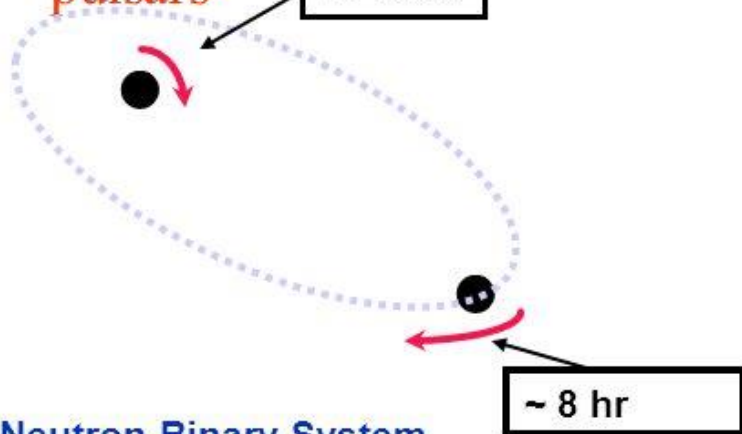
Delay in the
Time of Periastron
Of Binary Pulsar

Gravitational Waves

the evidence

Neutron Binary System – Hulse & Taylor

PSR 1913 + 16 -- Timing of pulsars



Neutron Binary System

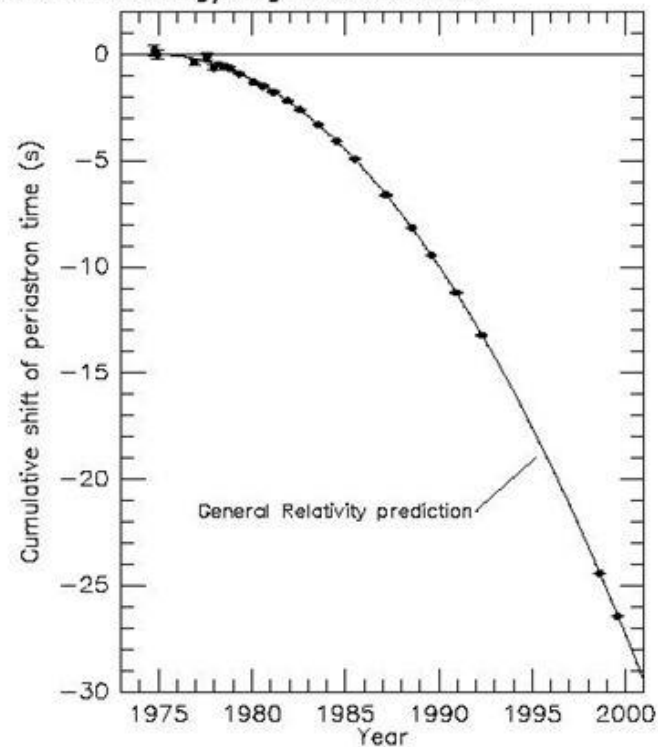
- separated by 10^6 miles
- $m_1 = 1.4m_{\odot}$; $m_2 = 1.36m_{\odot}$; $\varepsilon = 0.617$

Prediction from general relativity

- spiral in by 3 mm/orbit
- rate of change orbital period

Emission of gravitational waves

Comparison between observations of the binary pulsar PSR1913+16, and the prediction of general relativity based on loss of orbital energy via gravitational waves

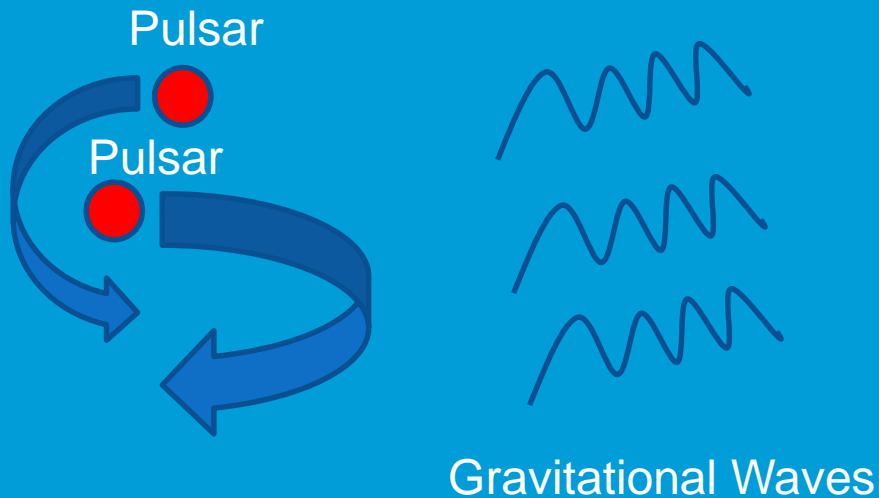


From J. H. Taylor and J. M. Weisberg, unpublished (2000)

5. Direct Detection

Two massive and compact objects orbiting one around the other

- Gravity as a fundamental force
- General Relativity
- Radiation of Gravitational Waves
- Indirect evidence (1974)
- **Direct Detection (2016)**



Detector on Earth



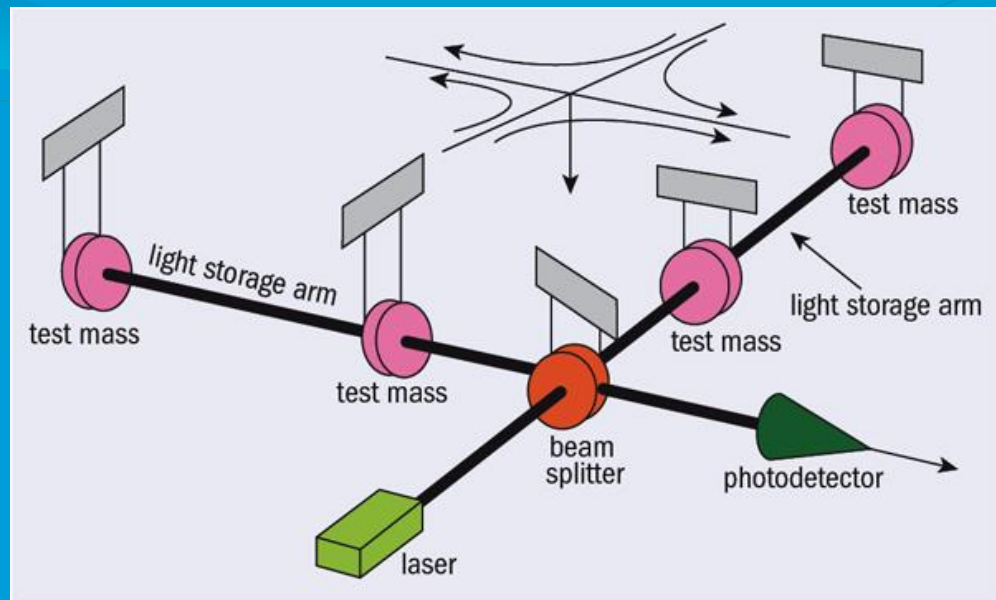
1960: J. Weber proposal of resonant detectors (cryogenic bars, 1960).

In Europe three of them have been built by E. Amaldi and coworkers (CERN, Legnaro and Frascati).

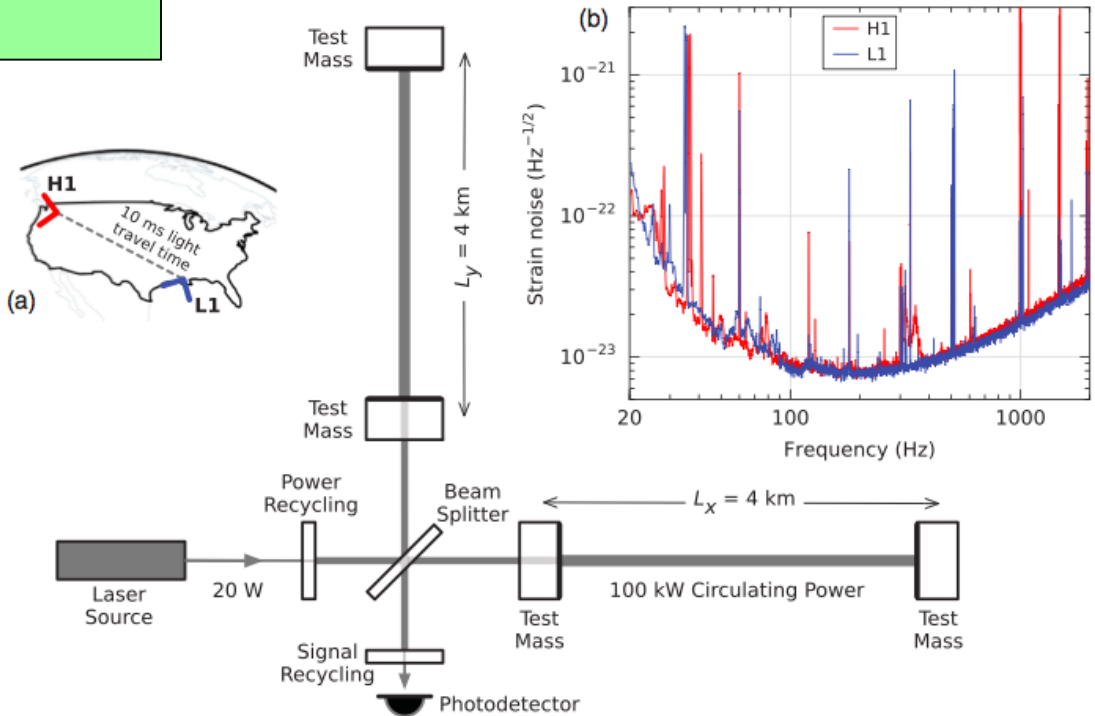
The search for Gravitational Waves: Michelson Interferometers in the Fabry-Pérot configuration

Key idea: mirrors (test masses)
whose position is monitored by laser

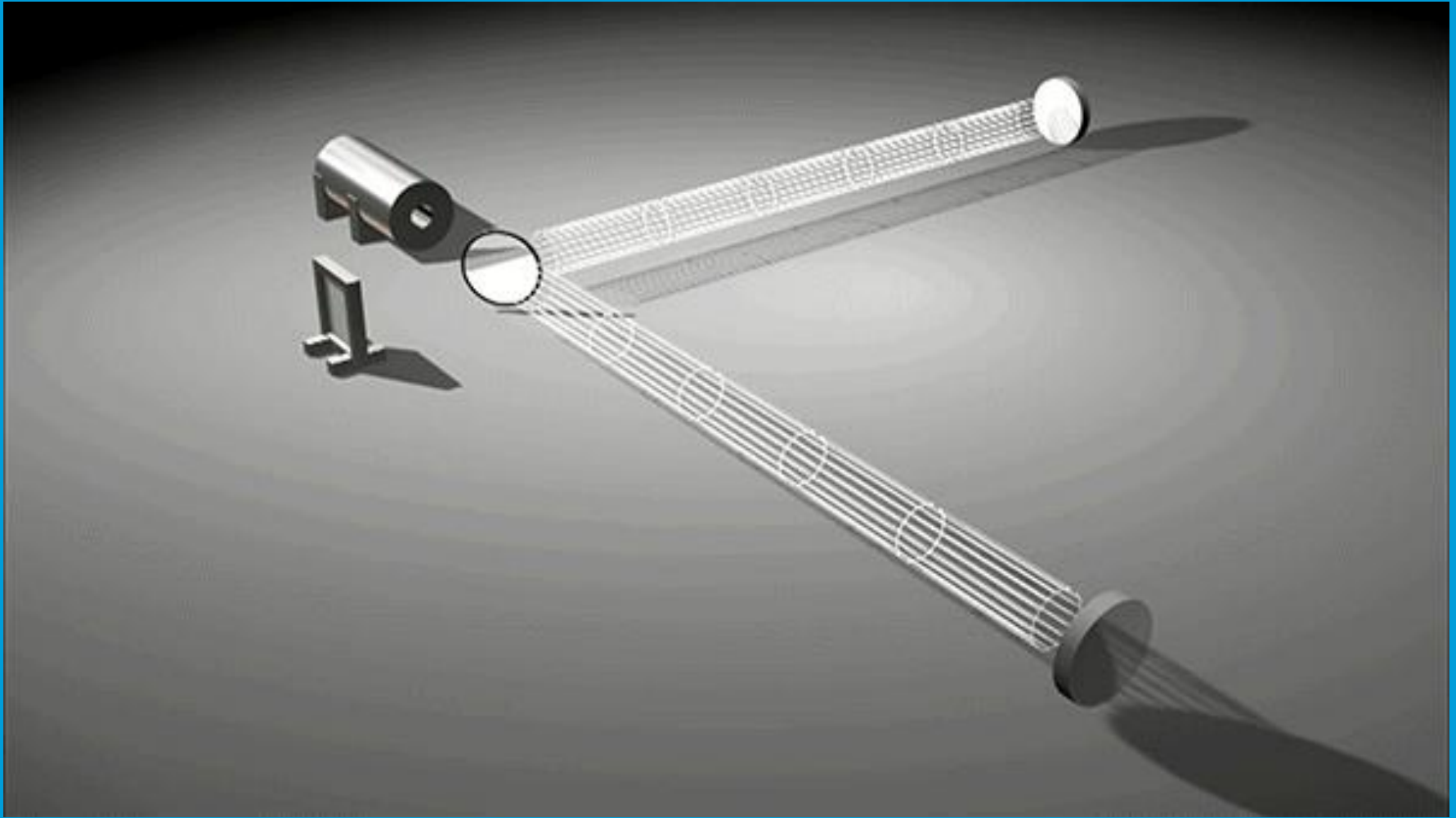
LIGO (USA, 2 interferometers) and
Virgo (Italy-France) in a single 1000
scientist collaboration



Strain Sensitivity and
location of Hanford and
Livingston Interferometers
(LIGO)



The search for Gravitational Waves: Michelson Interferometer to detect the «mirror position during the passage of the gravitational wave».



A gravitational wave will actually «create and destroy» space.

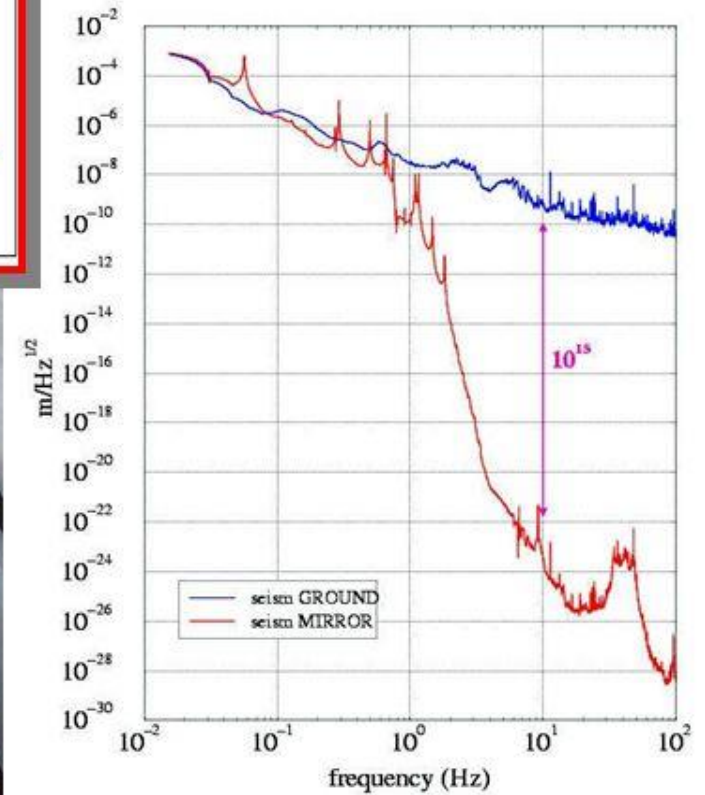
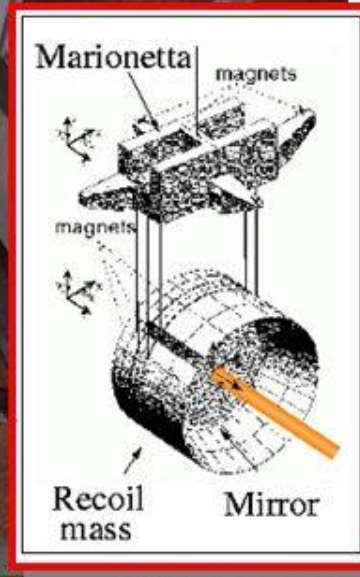
Inside one of the LIGO arms



Superattenuators

Possible contributions:

- Virgo+ will use monolithic suspension
- Input-mode cleaner suspension



14 September 2016: Hanford and Livingston observe at the same time (within 10 ms) a clear gravitational wave signal: GW150914, with full duration of 0.5 s.

This kind of signal can be generated only by the mutual collapse of two Black-Holes having ~36 and 29 Solar Masses. The resulting (Kerr-type) Black Hole has 62 solar masses. Three solar masses have been transformed to energy (spacetime waves).

Physical Review Letters 116 (2016) 061102.

Raw Signals

Background subtracted

Background

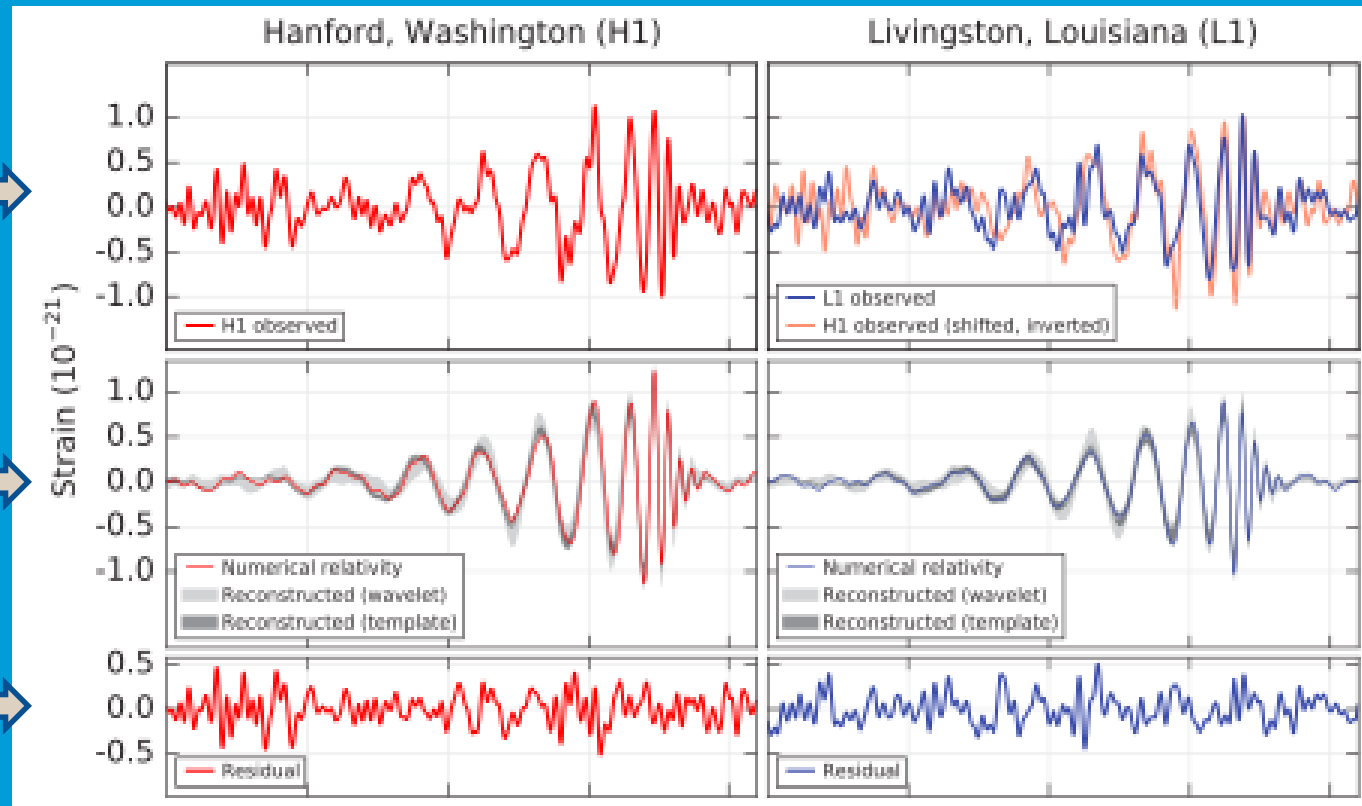
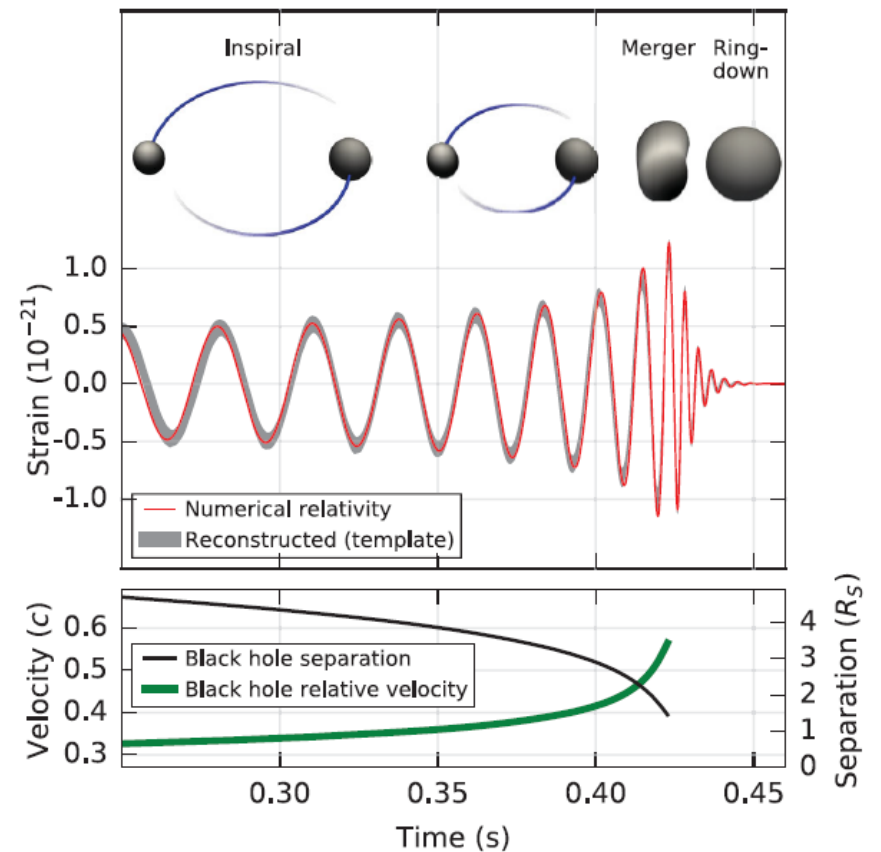


TABLE I. Source parameters for GW150914. We report median values with 90% credible intervals that include statistical errors, and systematic errors from averaging the results of different waveform models. Masses are given in the source frame; to convert to the detector frame multiply by $(1+z)$ [90]. The source redshift assumes standard cosmology [91].

| | |
|---------------------------|--------------------------|
| Primary black hole mass | $36^{+5}_{-4} M_{\odot}$ |
| Secondary black hole mass | $29^{+4}_{-4} M_{\odot}$ |
| Final black hole mass | $62^{+4}_{-4} M_{\odot}$ |
| Final black hole spin | $0.67^{+0.05}_{-0.07}$ |
| Luminosity distance | 410^{+160}_{-180} Mpc |
| Source redshift z | $0.09^{+0.03}_{-0.04}$ |



The evidence in favor of a BH-BH scenario is compelling (only two BH's can reach 75 Hz of mutual rotational frequency before merging in a region of

$$R_S \gg 200 \text{ km}$$

Maximum power transformed into gravitational waves is about 1000 times a Supernova energy release!

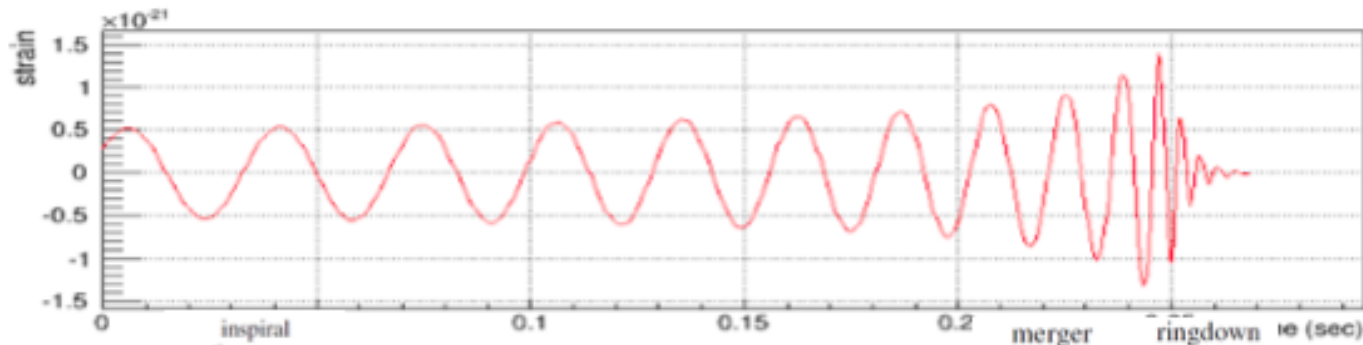
$$P_{\text{max}} \approx 4 \times 10^{56} \text{ erg/s}$$

Spectacular confirmation of General Relativity as a classical theory.

Well predicted discovery (this slide from a 2014 Conference)!

Sources of gravitational waves

- Compact binary (BNS, BBH) coalescence. Best candidate for ground based detectors



The GW signal produced in the last few inspiraling cycles are expected to fall in the interferometer bandwidth

GW emission well approximated by the quadrupole formula. Analytical solution available. GW Candle.



Only numerical solution available

Perturbative or numerical

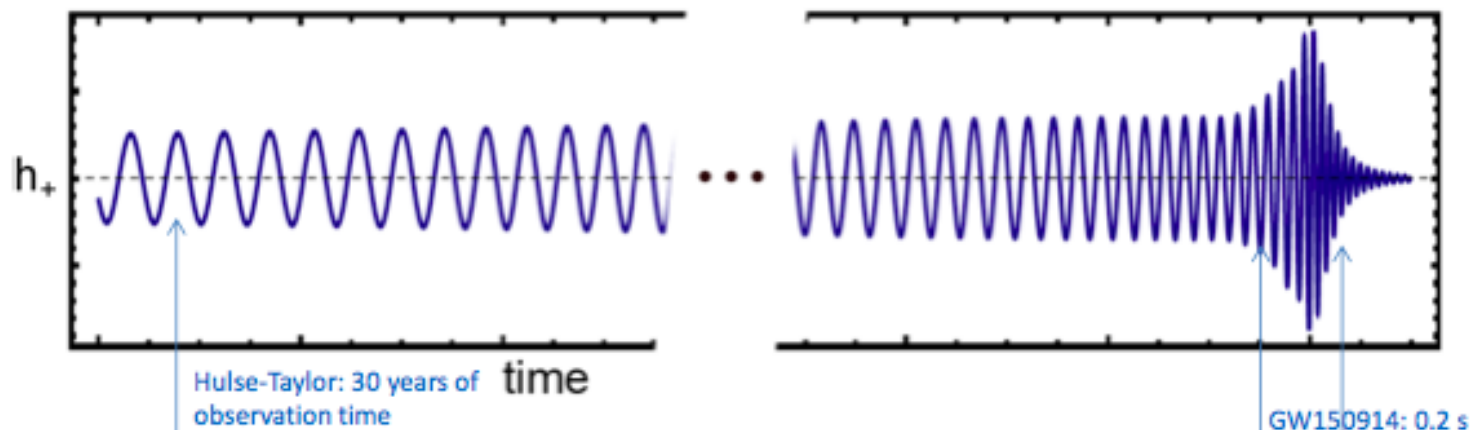
Test of general relativity in strong non-linear regime

Numerical (not Exact)
General Relativity Solutions

How does GW150914 compare with the Hulse-Taylor binary pulsar?

PSR1916+13 versus GW150914

Waveform



| | | |
|---|------------------------|---------------------------------------|
| PSR1916+13 | Binary system | GW150914 |
| NS-NS | Compact object | BH-BH |
| $M_1 = 1.44 M_\odot, M_2 = 1.3 M_\odot$ | Mass | $M_1 = 36 M_\odot, M_2 = 29 M_\odot$ |
| 4×10^{-23} | GW amplitude | 2×10^{-21} |
| $7 \times 10^{-5} \text{ Hz}$ | GW frequency | $30 \div 300 \text{ Hz}$ |
| $7.4 \times 10^9 \text{ years}$ | Time to merging | $0.3 \div 0 \text{ s}$ |
| $6 \times 10^{30} \text{ erg s}^{-1}$ | Peak luminosity | $3 \times 10^{56} \text{ erg s}^{-1}$ |
| 6.4 pc | Distance | 410 Mpc |
| 10^6 km | Radius orbit | $\sim 200 \text{ km}$ |

Detection of other Gravitational Waves sources? Yes!

PRL 116, 241103 (2016)

PHYSICAL REVIEW LETTERS

week ending
17 JUNE 2016

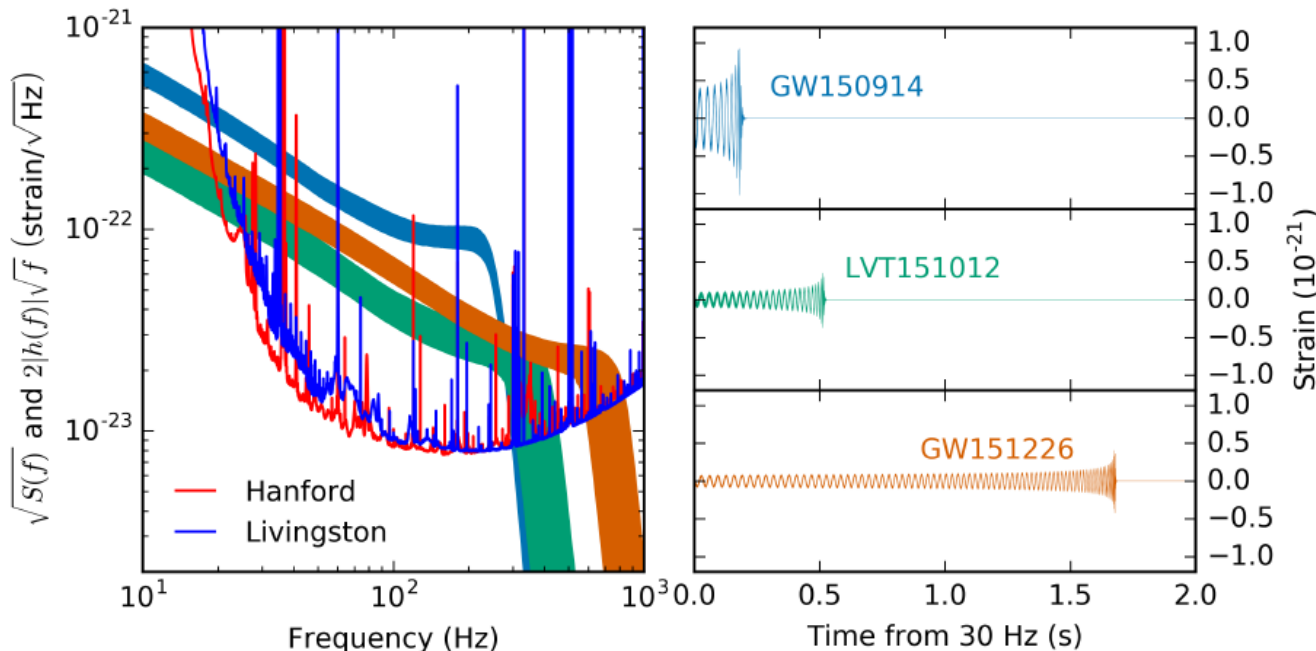


GW151226: Observation of Gravitational Waves from a 22-Solar-Mass Binary Black Hole Coalescence

B. P. Abbott *et al.**

(LIGO Scientific Collaboration and Virgo Collaboration)

(Received 31 May 2016; published 15 June 2016)



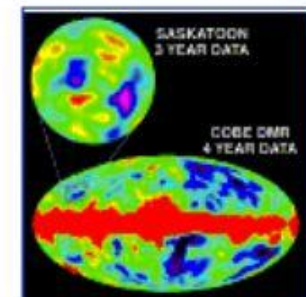
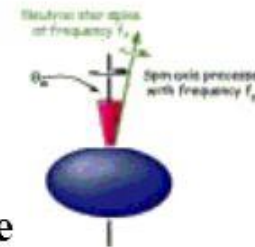
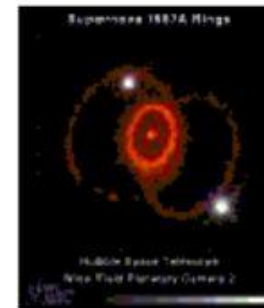
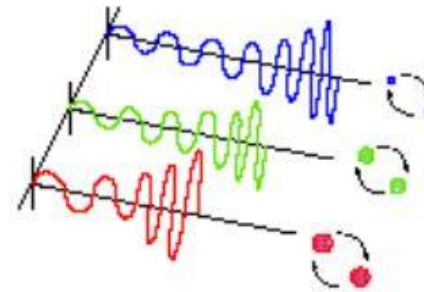
Second direct observation published in June 2016!

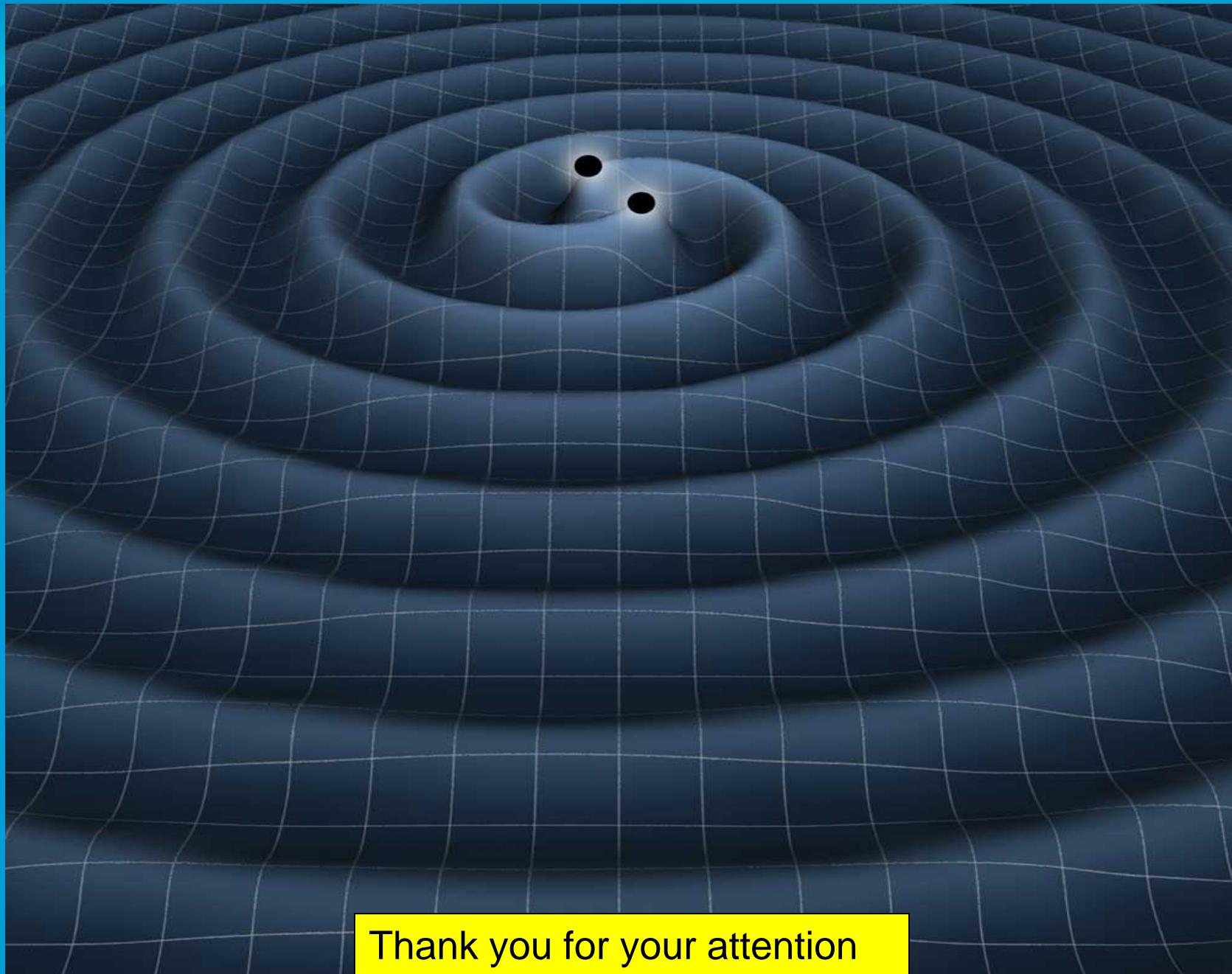
Observed Black Hole – Black Hole (BH-BH) coalescences are at least two. Maybe three.

The expectation is for many sources to be identified in the near future. A Gravitational Wave astronomy!

Astrophysical Sources for Terrestrial GW Detectors

- **Compact binary Coalescence: “chirps”**
 - NS-NS, NS-BH, BH-BH
- **Supernovas or GRBs: “bursts”**
 - GW signals observed in coincidence with EM or neutrino detectors
- **Pulsars in our galaxy: “periodic waves”**
 - Rapidly rotating neutron stars
 - Modes of NS vibration
- **Cosmological: “stochastic background”?**
 - Probe back to the Planck time (10^{-43} s)
 - Probe phase transitions : window to force unification
 - Cosmological distribution of Primordial black holes

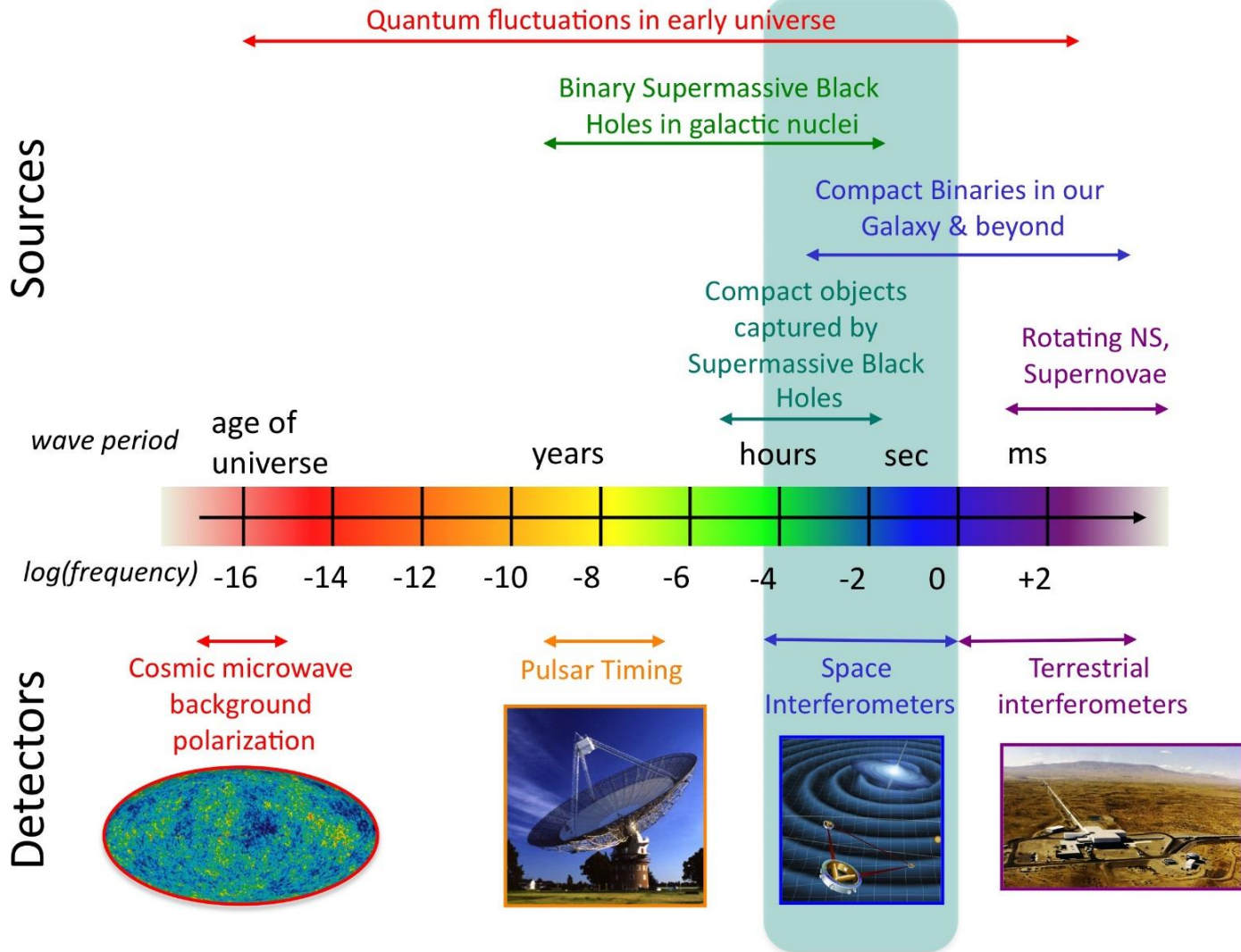




Thank you for your attention

Backup Slides

The Gravitational Wave Spectrum

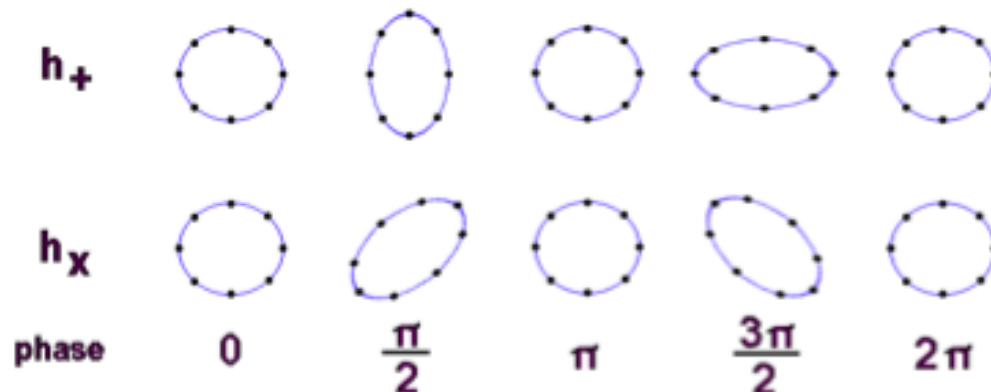


Gravitational Waves and G.R. (1916)

- Gravity is a manifestation of curvature of space-time produced by matter-energy
- Any rapidly moving mass generates fluctuations in spacetime curvature which propagate at the speed of light. **Gravitational waves**
- The physical quantity transported by gravitational waves is curvature

GW properties

- Speed of light,
- Transverse, traceless,
- GWs stress and compress spacetime in two directions,
- Two polarizations states “+” and “x”.



$$\frac{\Delta L}{L} \sim h$$

Sources of gravitational waves

A mass distribution with $\ddot{Q}_{\mu\nu} \neq 0$ ($Q_{\mu\nu}$ mass-quadrupole) produces GWs

$$P = \frac{G}{5c^5} \ddot{Q}_{\mu\nu} \ddot{Q}^{\mu\nu} \quad \text{Even in the most optimistic case } h \leq 10^{-20}$$

Very low

- ❑ Burst (i.e. core collapse supernovae)

$$h \sim 6 \times 10^{-21} \left(\frac{E}{10^{-7} M_{\odot} c^2} \right)^{\frac{1}{2}} \left(\frac{1 \text{ ms}}{T} \right) \left(\frac{1 \text{ kHz}}{f} \right) \left(\frac{10 \text{ kpc}}{r} \right)$$

- ❑ Continuous waves (i.e non axisymmetric spinning neutron star)

$$h_0 = \frac{4\pi^2 G I_{zz} f_{GW}^2}{c^4 r} \epsilon = (1.1 \times 10^{-24}) \left(\frac{I_{zz}}{I_0} \right) \left(\frac{f_{GW}}{1 \text{ kHz}} \right)^2 \left(\frac{1 \text{ kpc}}{r} \right) \left(\frac{\epsilon}{10^{-6}} \right) \quad \epsilon \equiv \frac{I_{xx} - I_{yy}}{I_{zz}}$$

- ❑ Stochastic both from cosmological or astrophysical origin

$$[S_{GW}(f)]^{\frac{1}{2}} = (5.6 \times 10^{-22}) h_{100} (\Omega(f))^{\frac{1}{2}} (100 \text{ Hz})^{\frac{3}{2}} \text{ Hz}^{-\frac{1}{2}}$$


GWs: more properties

- EM waves: at lowest order the radiation can be emitted by a dipole source ($l=1$). Monopolar radiation is forbidden as a result of charge conservation.
- GWs: the lowest allowed multipole is the **quadrupole** ($l=2$). The monopole is forbidden as a result of mass conservation. Similarly, dipole radiation is absent as a result of momentum conservation.
- GWs represents propagating “ripples in spacetime” or, more accurately, a **propagating curvature perturbation**. The perturbed curvature (Riemann tensor) is given by (in the TT gauge):

$$R_{j0k0}^{\text{TT}} = -\frac{1}{2} \partial_t^2 h_{jk}^{\text{TT}}, \quad j, k = 1, 2, 3$$

GWs and curvature

- As we mentioned, GWs represent a fluctuating curvature field.
- Their effect on test particles is of tidal nature.
- Equation of **geodesic deviation** (in weak gravity):

$$\frac{d^2 \xi^k}{dt^2} = -R_{0j0}^k{}^{\text{TT}} \xi^j$$


distance between geodesics (test particles)

- Newtonian limit: $R_{k0j0}^{\text{TT}} \approx \frac{\partial^2 \Phi}{\partial x^k \partial x^j}$ ← Newtonian grav. potential

GWs vs EM waves

- Similarities:

- ✓ Propagation with the speed of light.
- ✓ Amplitude decreases as $\sim 1/r$.
- ✓ Frequency redshift (Doppler, gravitational, cosmological).

- Differences:

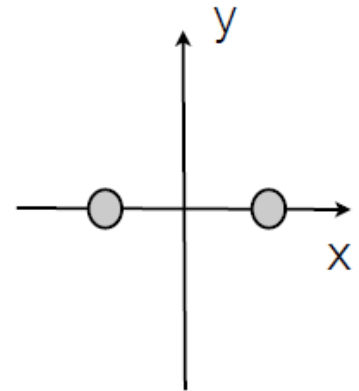
- ✓ GWs propagate through matter with little interaction. Hard to detect, but they carry uncontaminated information about their sources.
- ✓ Strong GWs are generated by bulk (coherent) motion. They require strong gravity/high velocities (compact objects like black holes and neutron star).
- ✓ EM waves originate from small-scale, incoherent motion of charged particles. They are subject to “environmental” contamination (interstellar absorption etc.).

Effect on test particles (I)

- We consider a pair of test particles on the cartesian axis Ox at distances $\pm x_0$ from the origin and we assume a GW traveling in the z -direction.
- Their distance will be given by the relation:

$$\begin{aligned} dl^2 &= g_{\mu\nu} dx^\mu dx^\nu = \dots = -g_{11} dx^2 = \\ &= (1 - h_{11})(2x_0)^2 = [1 - h_+ \cos(\omega t)] (2x_0)^2 \end{aligned}$$

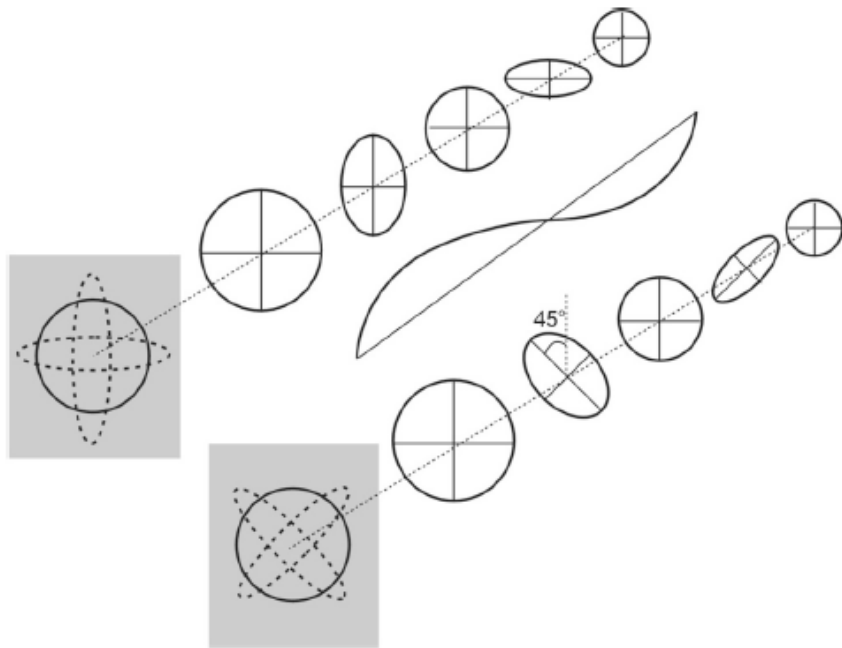
$$dl \approx \left[1 - \frac{1}{2} h_+ \cos(\omega t) \right] (2x_0)$$



Effect on test particles (II)

- Similarly for a pair of particles placed on the y-axis:

- Comment: the same result can be derived using the geodetic deviation equation. $dl \approx \left[1 + \frac{1}{2} h_+ \cos(\omega t) \right] (2y_0)$



The quadrupole formula

- Einstein (1918) derived the quadrupole formula for gravitational radiation by solving the linearized field equations with a source term:

$$\square h^{\mu\nu}(t, \vec{x}) = -\kappa T^{\mu\nu}(t, \vec{x}) \longrightarrow h^{\mu\nu} = -\frac{\kappa}{4\pi} \int_V d^3x' \frac{T^{\mu\nu}(t - |\vec{x} - \vec{x}'|, \vec{x}')}{|\vec{x} - \vec{x}'|}$$

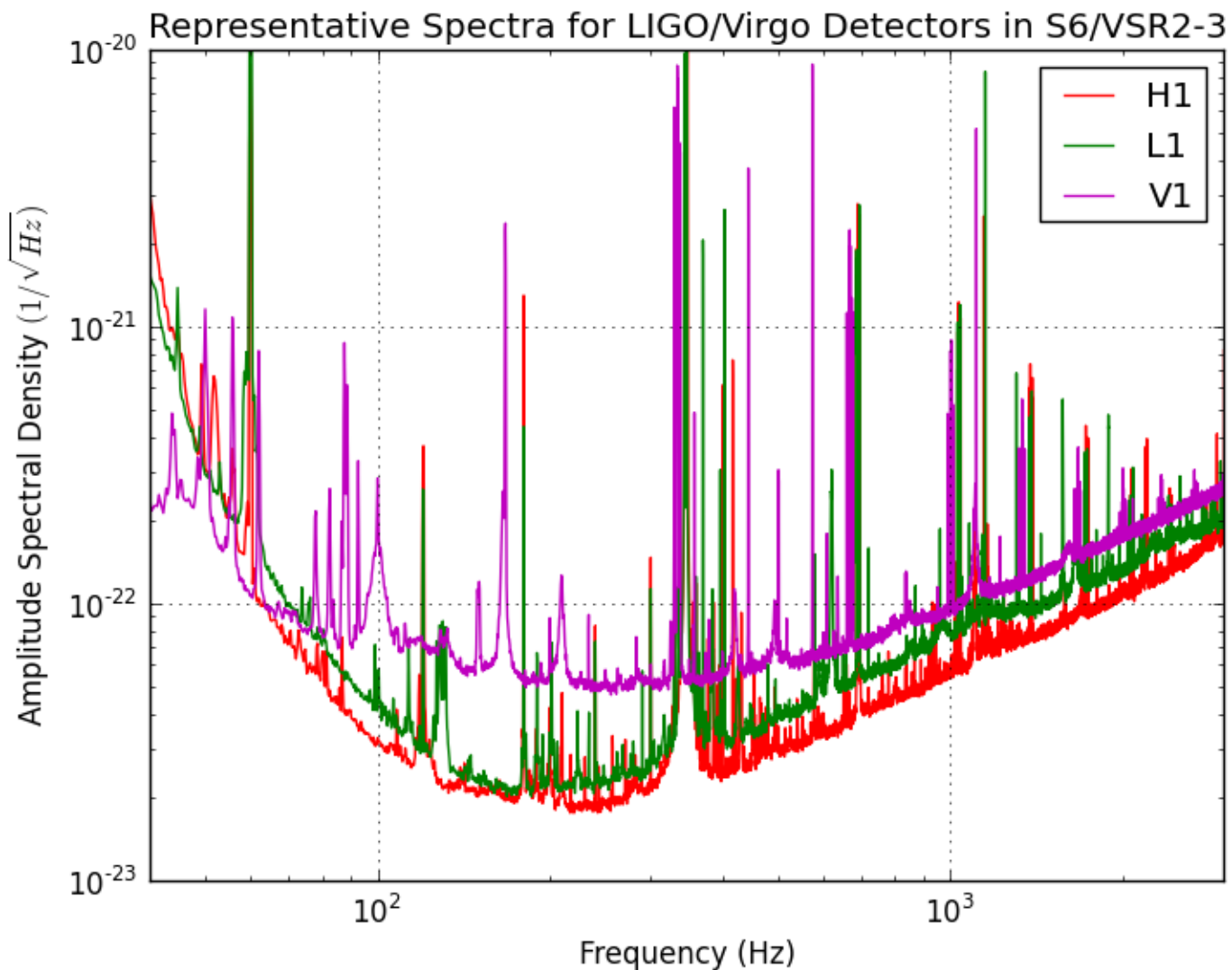
- This solution suggests that the wave amplitude is proportional to the **second time derivative of the quadrupole** moment of the source:

$$h^{\mu\nu} = \frac{2G}{r c^4} \ddot{Q}_{\text{TT}}^{\mu\nu}(t - r/c) \qquad Q_{\text{TT}}^{\mu\nu} = \int d^3x \rho \left(x^\mu x^\nu - \frac{1}{3} \delta^{\mu\nu} r^2 \right)$$

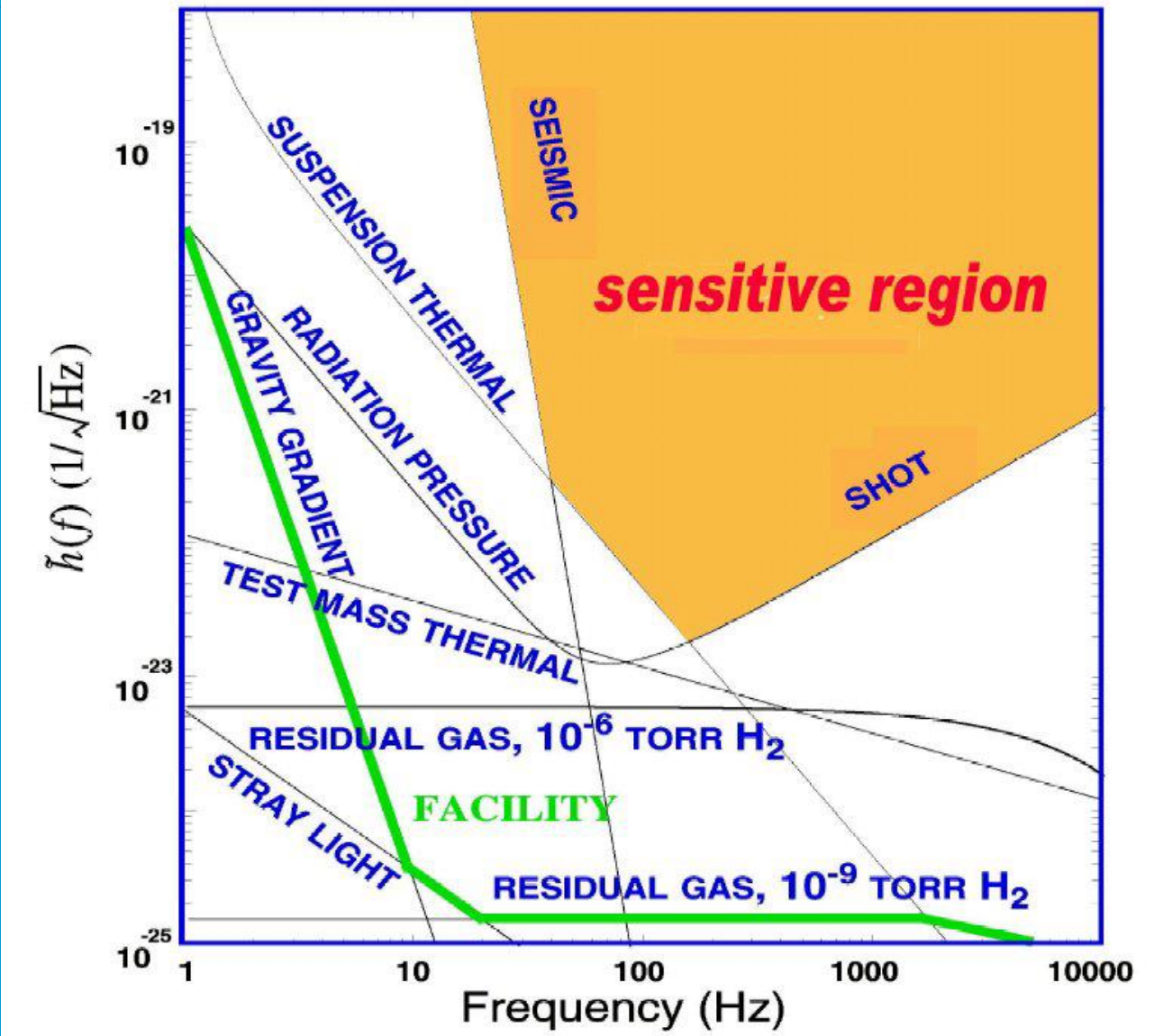
(quadrupole moment in the “TT gauge” and at the retarded time $t-r/c$)

- This result is quite accurate for all sources, as long as the wavelength is much longer than the source size R .

Sensitività degli osservatori Hanford (LIGO), Livingston (LIGO), Cascina (Virgo)
nel 2012 circa

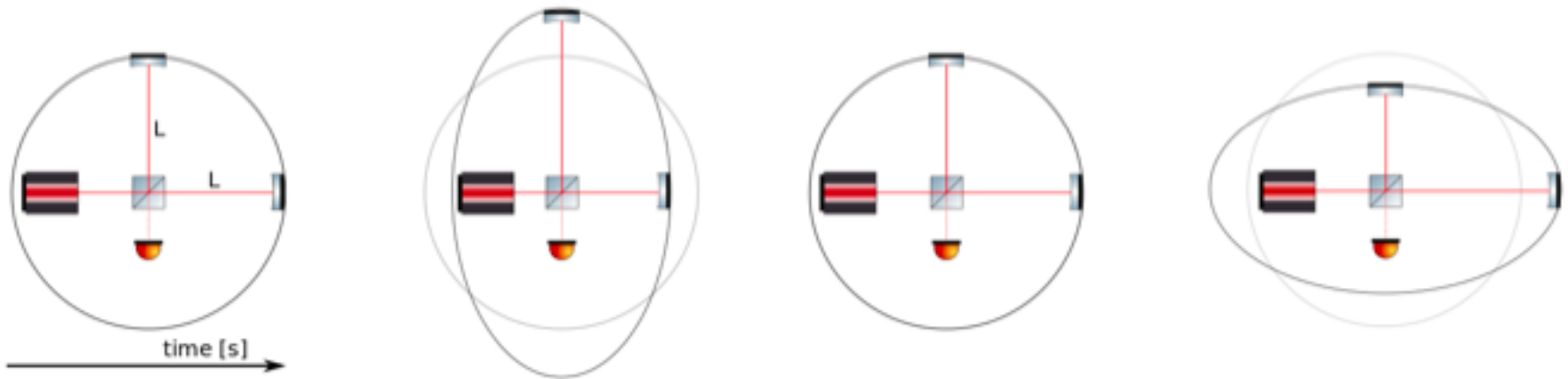


Noise structure in gravitational interferometers



Interferometers

Interferometer response to h_+



$$\Delta\phi = 4\pi \frac{Lh_+}{\lambda_{Laser}} \frac{\sin\left(\frac{\Omega_{GW}L}{c}\right)}{\frac{\Omega_{GW}L}{c}}$$

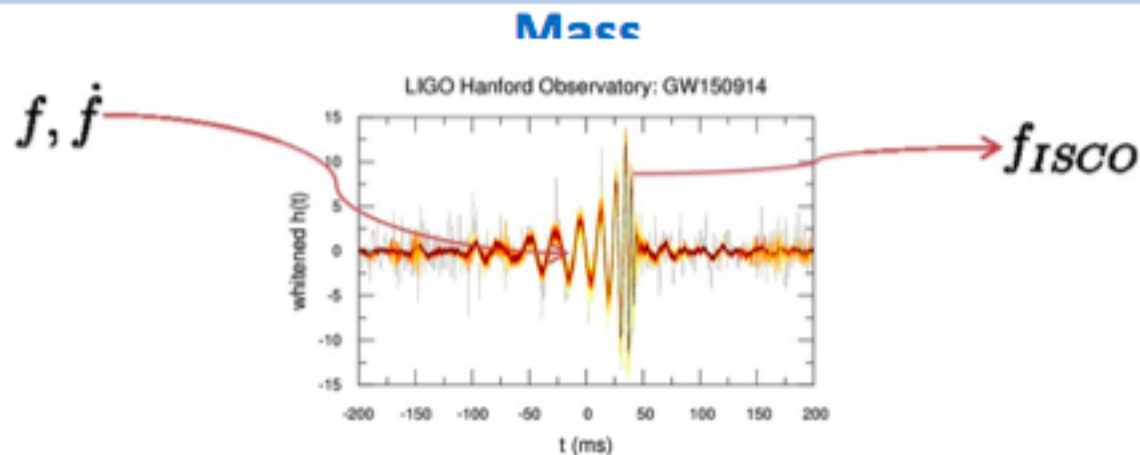
- Broad band
- Sensitivity increases with L
- Cutoff frequency:

$$\frac{\Omega_{GW}L}{c} \sim \frac{\pi}{2} \Rightarrow L < \frac{\lambda_{GW}}{4}$$

Anisotropic response and polarization sensitivity

$$h = F_{(+)}h_+ + F_{(\times)}h_\times \quad \left\{ \begin{array}{l} h \quad \text{strain measured} \\ F_{(+,\times)} \quad \text{(sky position, polarization)} \end{array} \right.$$

GW150914 The binary parameters



Rough estimation: Leading order in the PN parameters and circular orbits

- Observed GW frequency f and time derivative \dot{f} gives the «chirp mass» M_c

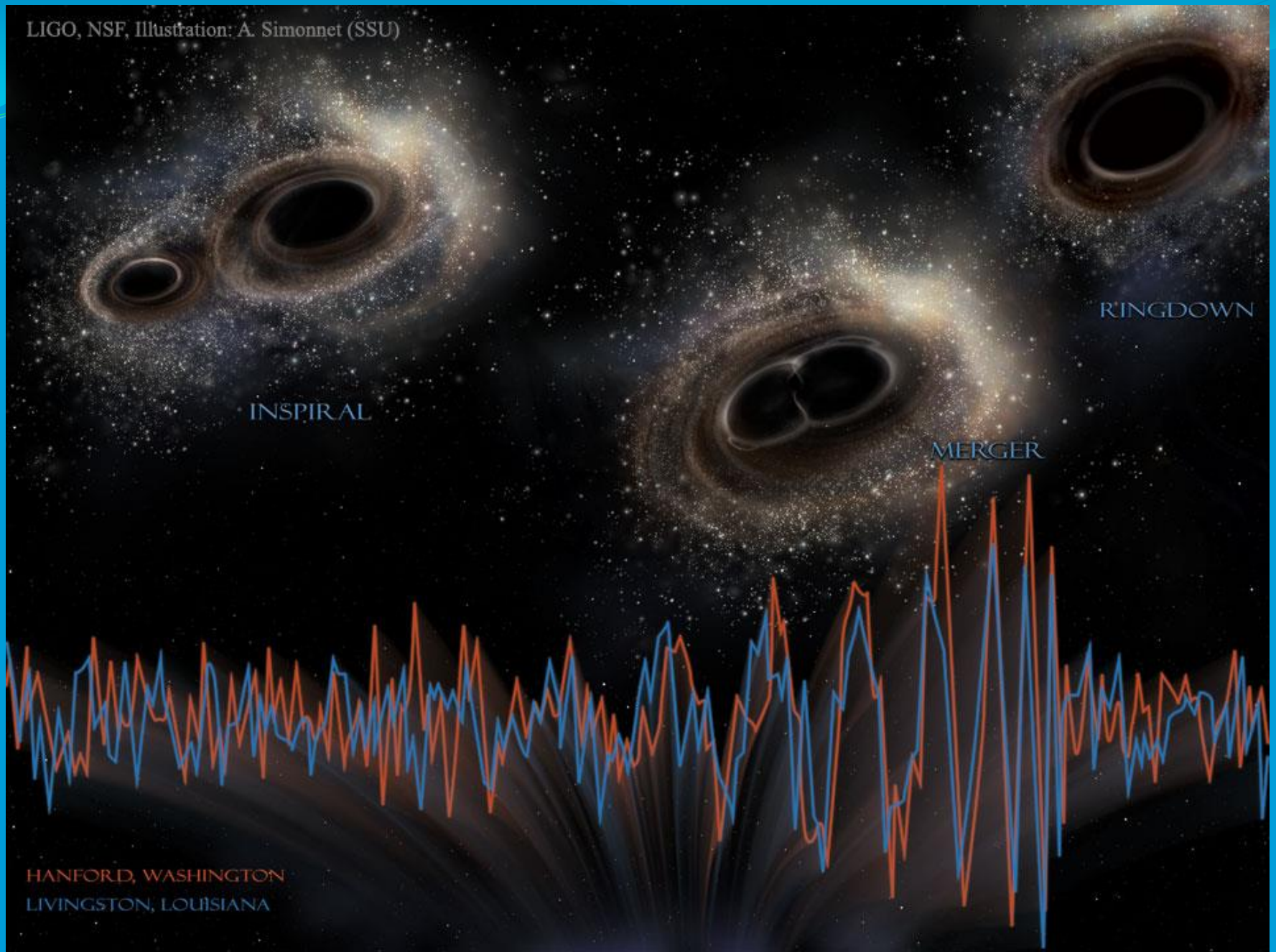
$$M_c = \frac{(m_1 m_2)^{3/5}}{(m_1 + m_2)^{1/5}} = \frac{c^3}{G} \left[\frac{5}{96} \pi^{-3/8} f^{-11/3} \dot{f} \right]^{3/4} \approx 30 M_\odot$$

- Frequency of the Schwarzschild innermost stable circular orbit (ISCO) gives the total mass

$$f_{ISCO} \approx 4400 \frac{M_\odot}{M} [\text{Hz}]$$

$$f_{ISCO} \approx 150 \text{ Hz} \Rightarrow M_{Tot} = m_1 + m_2 \approx 70 M_\odot$$

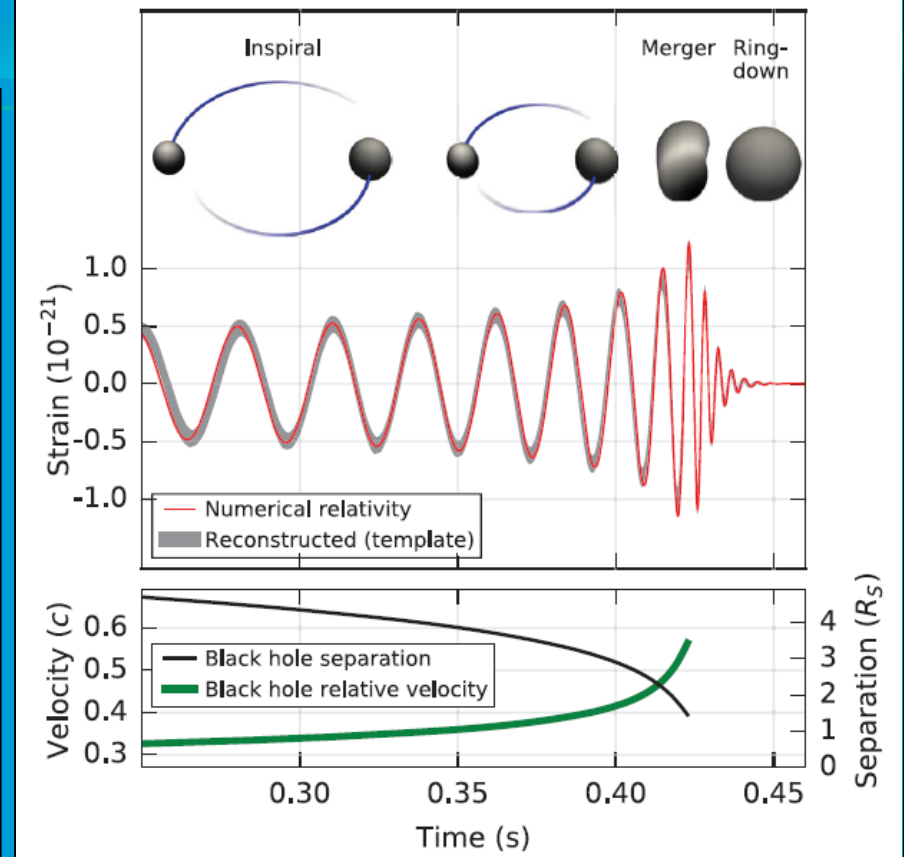
$$m_1 \approx m_2 \approx 30 M_\odot$$



High-Energy Events in the Universe

TABLE I. Source parameters for GW150914. We report median values with 90% credible intervals that include statistical errors, and systematic errors from averaging the results of different waveform models. Masses are given in the source frame; to convert to the detector frame multiply by $(1+z)$ [90]. The source redshift assumes standard cosmology [91].

| | |
|---------------------------|---------------------------------|
| Primary black hole mass | $36_{-4}^{+5} M_{\odot}$ |
| Secondary black hole mass | $29_{-4}^{+4} M_{\odot}$ |
| Final black hole mass | $62_{-4}^{+4} M_{\odot}$ |
| Final black hole spin | $0.67_{-0.07}^{+0.05}$ |
| Luminosity distance | $410_{-180}^{+160} \text{ Mpc}$ |
| Source redshift z | $0.09_{-0.04}^{+0.03}$ |



Maximum power transformed into gravitational waves is about 1000 times a Supernova energy release!

$$P_{\max} \approx 4 \times 10^{56} \text{ erg/s}$$

Compare with Planck power

$$W_P = \frac{E_P}{t_P} = \sqrt{\frac{\hbar c^5}{2G}} \sqrt{\frac{c^5}{G\hbar}} = \frac{c^5}{G\sqrt{2}} = 2.5 \times 10^{59} \text{ erg s}^{-1}$$

A Supernova

Total energy

$$E \approx 0.05 M_{sun} c^2 = 10^{53} \text{ erg}$$

90% of energy goes into neutrinos in 10 s

Avg Power

$$P_{avg} \approx 10^{52} \text{ erg / s}$$

A Gamma-Ray Burst

Total energy

$$E \approx 0.5 M_{sun} c^2 = 10^{54} \text{ erg}$$

90% of energy goes into e.m. waves or gamma rays in 20 s

Avg Power

$$P_{avg} \approx 5 \times 10^{52} \text{ erg / s}$$

GW150914

Total energy

$$E \approx 3 M_{sun} c^2 = 10^{55} \text{ erg}$$

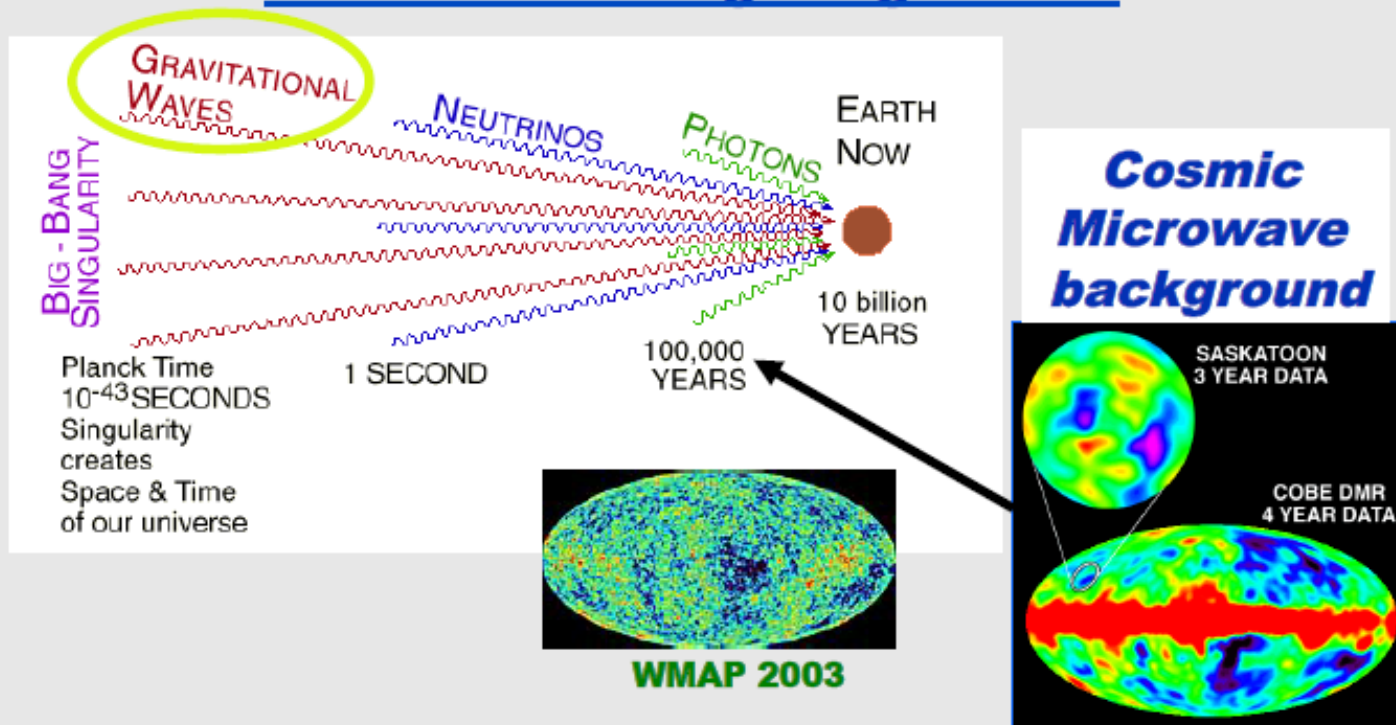
All energy goes into gravitational waves in 0.5 s

Avg Power

$$P_{avg} \approx 5 \times 10^{55} \text{ erg / s}$$

Universo Primordiale: *“rumore correlato tra i vari rivelatori”*

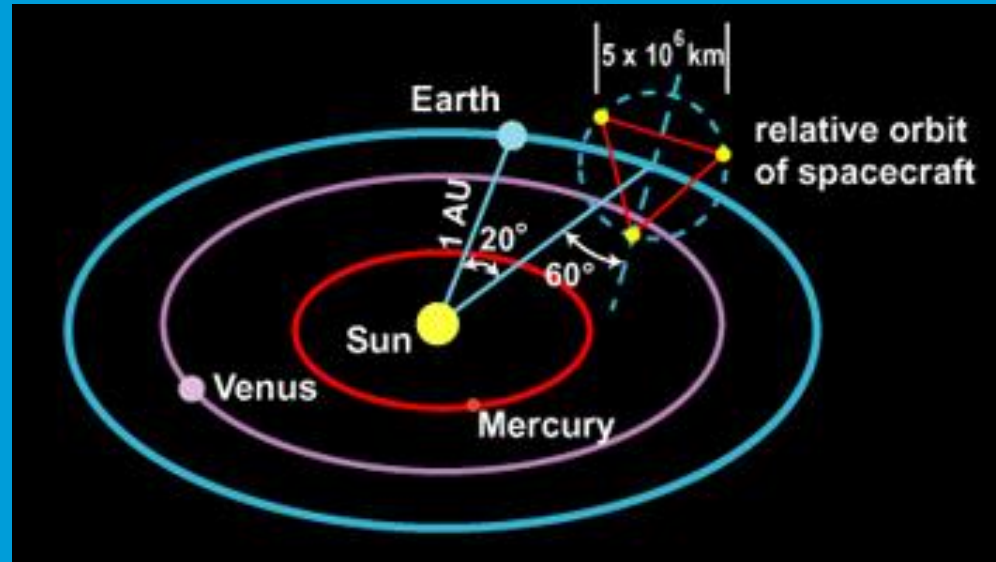
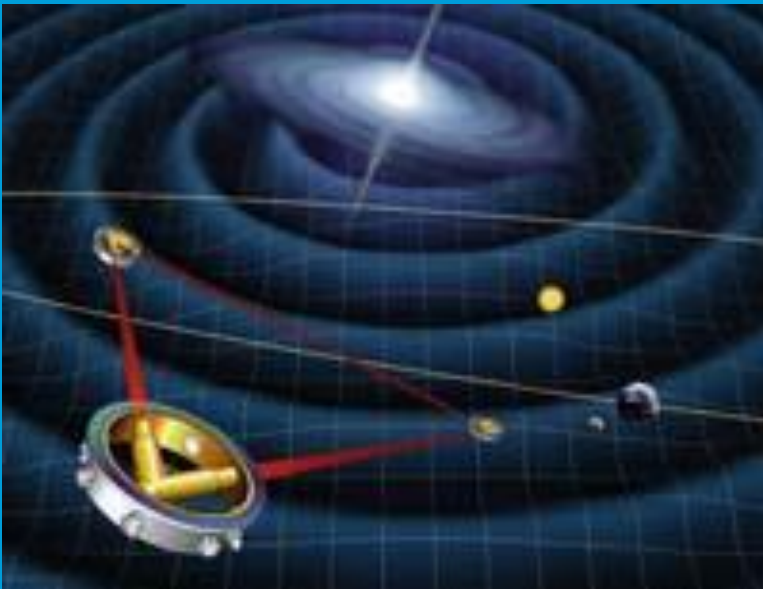
‘Mormorio’ dal Big Bang iniziale



27

Space-based GW detection

- LISA (*Laser Interferometer Space Antenna*)



Four-Velocity and Four-Momentum

Four-velocity: a vector tangent to the world line of the particle, and stretching a unit time in the particle frame:

Let us consider a uniformly moving particle: in its rest frame the four-velocity points along the time axis and it is a unit time length

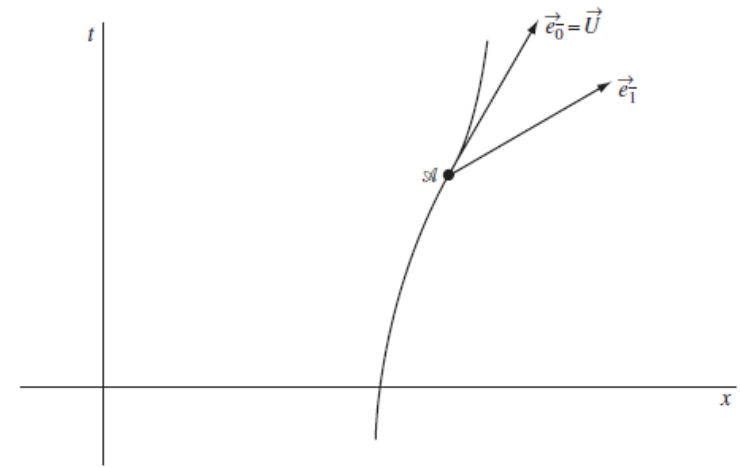
More in general: The four-velocity is the time basis vector of a particle in the MCRF

Let us define the four-momentum as

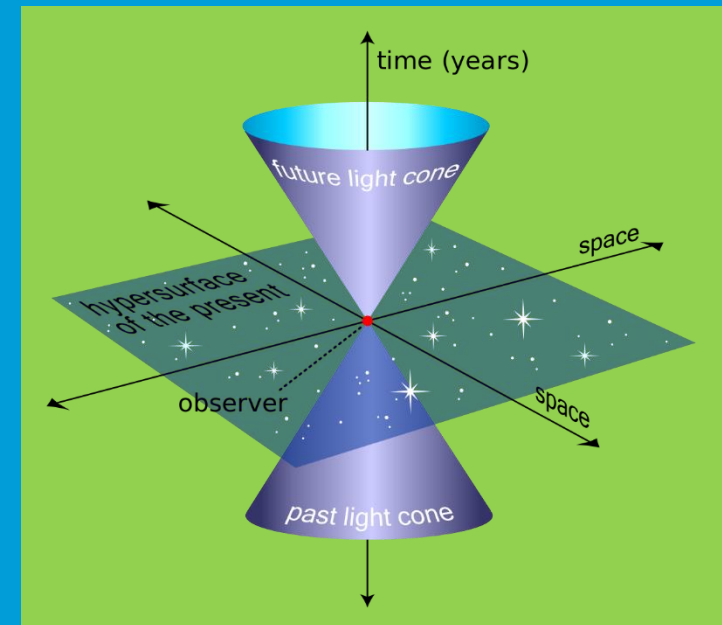
$$\vec{p} = m\vec{U}$$

And call its components in an inertial frame O

$$\vec{p} \rightarrow {}_o(E, p^1, p^2, p^3)$$



The four-velocity and MCRF basis vectors of the world line at \mathcal{A}



Let us now consider a particle in motion in an inertial system O and transform its four-velocity from the rest system \bar{O} to the inertial system O (v along 1)

$$\vec{U} = \vec{e}_{\bar{0}} \quad \vec{p} = m\vec{e}_{\bar{0}}$$

$$\vec{U} \rightarrow_{\bar{O}} (1, 0, 0, 0)$$

$$\vec{p} \rightarrow_{\bar{O}} (m, 0, 0, 0)$$



$$U^\alpha = \Lambda_{\bar{\beta}}^\alpha (\vec{e}_{\bar{0}})^{\bar{\beta}} = \Lambda_{\bar{0}}^\alpha$$

$$p^\alpha = m\Lambda_{\bar{0}}^\alpha$$

So that in the O system we have

$$p^0 = mU^0 = m(1 - v^2)^{-1/2} = m\gamma$$

$$p^1 = mU^1 = mv(1 - v^2)^{-1/2} = mv\gamma$$

$$p^2 = p^3 = 0$$

General Relativity Math

Metrics $\vec{A}\vec{B} = A^\alpha B^\beta \vec{e}_\alpha \vec{e}_\beta = A^\alpha B^\beta \eta_{\alpha\beta}$ (4-vectors) $\eta_{\mu\nu} \rightarrow g_{\mu\nu}(x)$

$$ds^2 = dx^\alpha dx^\beta \eta_{\alpha\beta}$$

Curvilinear coordinates: Christoffel symbols (covariant derivative)

$$\frac{\partial \vec{V}}{\partial x^\beta} = \frac{\partial}{\partial x^\beta} (V^\alpha \vec{e}_\alpha) = \frac{\partial V^\alpha}{\partial x^\beta} \vec{e}_\alpha + V^\alpha \frac{\partial \vec{e}_\alpha}{\partial x^\beta} = V_{,\beta}^\alpha \vec{e}_\alpha + V^\alpha \Gamma_{\alpha\beta}^\mu \vec{e}_\mu = (V_{,\beta}^\alpha + V^\mu \Gamma_{\mu\beta}^\alpha) \vec{e}_\alpha = V_{;\beta}^\alpha \vec{e}_\alpha$$

$$V_{;\beta}^\alpha = V_{,\beta}^\alpha + V^\mu \Gamma_{\mu\beta}^\alpha \quad p_{\alpha;\beta} = p_{\alpha,\beta} - p_\mu \Gamma_{\alpha\beta}^\mu \quad g_{\alpha\beta;\mu} \equiv 0$$

Christoffel symbols and the metrics

$$\Gamma_{\beta\mu}^\gamma = \frac{1}{2} g^{\alpha\gamma} (g_{\alpha\beta,\mu} + g_{\alpha\mu,\beta} - g_{\beta\mu,\alpha})$$

Manifold: continuous space that locally looks (pseudo)Euclidean.
It is a set that can be continuously parametrized

Curvature can be introduced on the manifold

The surface of an infinite cylinder (or torus) is a manifold without curvature
The surface of a sphere is a manifold with curvature

We want only differentiable manifolds (a cone would not work, because of the apex)

Why introduce a metric over a differentiable manifold?

Because we need to express lengths and time intervals (and all physics)

A differentiable manifold on which a symmetric (0,2) tensor has been selected to act as the metrics is called a (pseudo)Riemannian manifold.

It is pseudo because the metrics is not positive-definite.

The metrics defines the curvature of the pseudo-Riemannian manifold.

Any symmetric matrix can be diagonalized preserving the number of positive and negative eigenvalues (the signature of the metrics).

Our capability to construct a local inertial frame:

$$g_{\mu\nu}(x) \rightarrow \eta_{\mu\nu} = \begin{bmatrix} -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix}$$

Local Flatness Theorem: at any point P on the manifold, a coordinate system can be found such that:

$$g_{\mu\nu}(P) = \eta_{\mu\nu}(P) + O[(x^\mu)^2]$$

(going to the local Lorentz frame). This is equivalent to:

$$g_{\mu\nu}(P) = \eta_{\mu\nu}$$

$$\frac{\partial}{\partial x^\gamma} g_{\mu\nu}(P) = 0$$

$$\frac{\partial^2}{\partial x^\gamma \partial x^\mu} g_{\mu\nu}(P) \neq 0$$

The curved space has a flat tangent space at any point
Straight lines in flat spacetimes are the world lines of free particles
So, free particles are moving on lines that are locally straight lines

$g_{\mu\nu, \gamma\mu}$

has 20 independent components

Lengths and Volumes :

$$l = \int |g_{\alpha\beta} dx^\alpha dx^\beta|^{1/2}$$

$$dx = (-g)^{1/2} dx' \quad g \equiv \det(g_{\alpha\beta})$$

Going to the Lorentz local frame in P:

$$g_{\mu\nu}(P) = \eta_{\mu\nu}$$

$$g_{\mu\nu,\gamma}(P) = 0$$

$$\Gamma_{\beta\mu}^{\gamma} = \frac{1}{2} g^{\alpha\gamma} (g_{\alpha\beta,\mu} + g_{\alpha\mu,\beta} - g_{\beta\mu,\alpha})$$

$$\Gamma_{\beta\mu}^{\gamma}(P) = 0$$

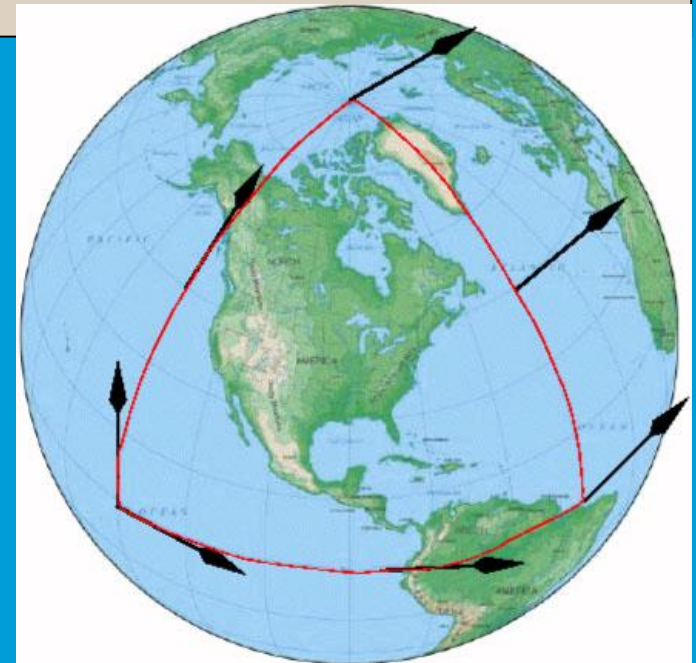
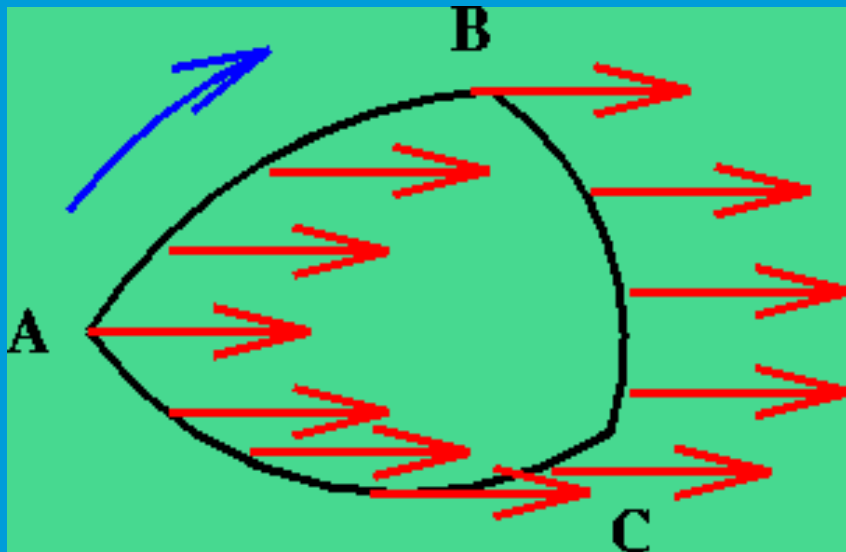
$$g_{\mu\nu,\gamma}(P) = g_{\mu\nu;\gamma}(P) = 0$$

$$g_{\mu\nu;\gamma} = 0$$

Now considering curvature :

- Extrinsic (just due to immersion in higher dimensional space). Euclid theorems holds and the space can be deformed (without tearing it or crumpling it) into flat space. Example: the cylinder.
- Intrinsic, like the surface of the sphere: parallel lines, when continued, do not remain parallel.

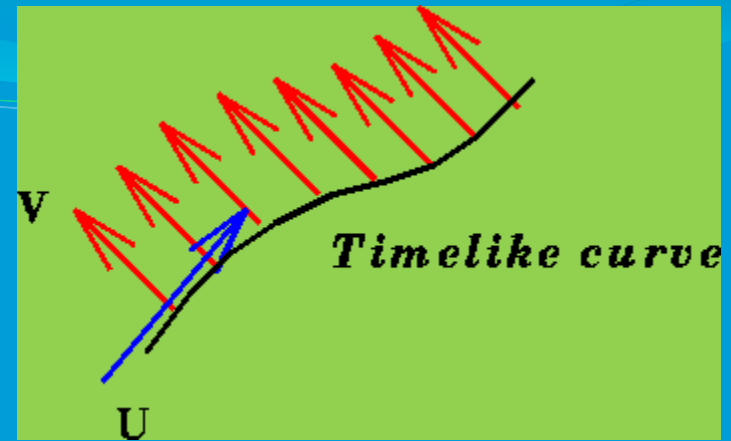
Parallel transport along a closed path: difference between a flat space (Euclidean plane) and a curved space (surface of the sphere).



Parallel transporting 4-vector V :

In a locally Lorentz frame :

$$\frac{dV^\alpha}{d\lambda} = 0 \rightarrow \frac{dV^\alpha}{dx^\beta} \frac{dx^\beta}{d\lambda} = V_{,\beta}^\alpha U^\beta = 0$$



The equation for parallel transport is then :

$$U^\beta V_{;\beta}^\alpha = 0 \Leftrightarrow \nabla_U \vec{V} = 0$$

Geodesics : a curve that parallel-transport its own tangent vector :

$$\nabla_U \vec{U} = 0$$

$$U^\beta U_{;\beta}^\alpha = U^\beta U_{,\beta}^\alpha + U^\beta U^\mu \Gamma_{\mu\beta}^\alpha = 0$$

If λ is a curve parameter this becomes :

$$U^\alpha = \frac{dx^\alpha}{d\lambda} \quad U^\beta \frac{\partial}{\partial x^\beta} = \frac{dx^\beta}{d\lambda} \frac{d}{dx^\beta} = \frac{d}{d\lambda}$$
$$\Rightarrow \frac{dU^\alpha}{d\lambda} + \Gamma_{\mu\beta}^\alpha \frac{dx^\beta}{d\lambda} \frac{dx^\mu}{d\lambda} = 0$$

$$\frac{d}{d\lambda} \frac{dx^\alpha}{d\lambda} + \Gamma_{\mu\beta}^\alpha \frac{dx^\beta}{d\lambda} \frac{dx^\mu}{d\lambda} = 0$$

Geodesic Equation. It is a curve of extremal length

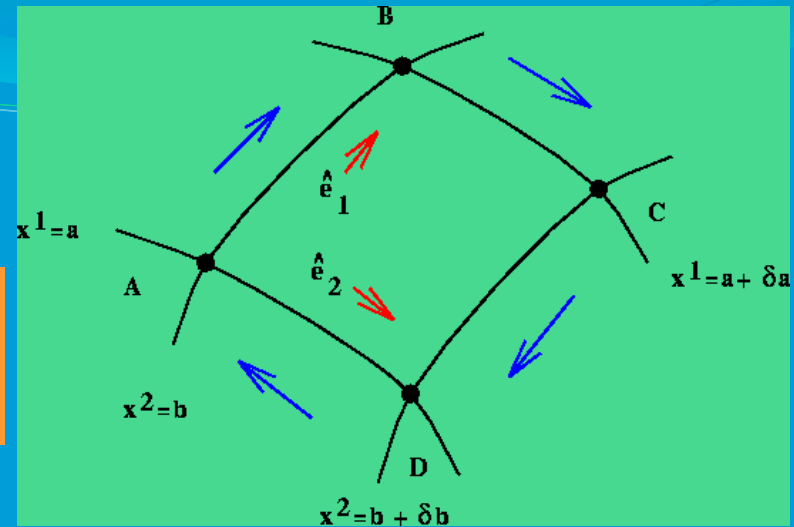
Curvature Tensor

Curvature: the effect of parallel transport around a closed loop

$$\delta V^\alpha = \delta a \delta b V^\mu \left[\Gamma_{\mu\sigma,\lambda}^\alpha - \Gamma_{\mu\lambda,\sigma}^\alpha + \Gamma_{\nu\lambda}^\alpha \Gamma_{\mu\sigma}^\nu - \Gamma_{\nu\sigma}^\alpha \Gamma_{\mu\lambda}^\nu \right] + \delta a \vec{e}_\sigma + \delta b \vec{e}_\lambda - \delta a \vec{e}_\sigma - \delta b \vec{e}_\lambda$$

$$\delta V^\alpha = \delta a \delta b V^\mu R_{\mu\lambda\sigma}^\alpha$$

$$R_{\beta\mu\nu}^\alpha = \Gamma_{\beta\nu,\mu}^\alpha - \Gamma_{\beta\mu,\nu}^\alpha + \Gamma_{\sigma\mu}^\alpha \Gamma_{\beta\nu}^\sigma - \Gamma_{\sigma\nu}^\alpha \Gamma_{\beta\mu}^\sigma$$



Riemann Curvature Tensor

What is the form of the Riemann Curvature Tensor in the local Lorentz frame P ?

$$\Gamma_{\beta\mu}^\gamma = \frac{1}{2} g^{\alpha\gamma} (g_{\alpha\beta,\mu} + g_{\alpha\mu,\beta} - g_{\beta\mu,\alpha}) \implies \Gamma_{\beta\mu}^\gamma (P) = 0 \implies g_{\mu\nu,\gamma} (P) = g_{\mu\nu;\gamma} (P) = 0$$

But the derivative of Γ is not zero

$$\Gamma_{\mu\nu,\sigma}^\alpha (P) = \frac{1}{2} g^{\alpha\beta} (g_{\beta\mu,\nu\sigma} + g_{\beta\nu,\mu\sigma} - g_{\mu\nu,\beta\sigma})$$

$$R_{\beta\mu\nu}^\alpha (P) = \frac{1}{2} g^{\alpha\sigma} (g_{\sigma\nu,\beta\mu} - g_{\sigma\mu,\beta\nu} + g_{\beta\mu,\sigma\nu} - g_{\beta\nu,\sigma\mu})$$

The following identities can be derived (which are true in P, but being tensor equations are valid everywhere)

$$R_{\alpha\beta\mu\nu} = -R_{\beta\alpha\mu\nu} = -R_{\alpha\beta\nu\mu} = R_{\mu\nu\alpha\beta}$$

$$R_{\alpha\beta\mu\nu} + R_{\alpha\nu\beta\mu} + R_{\alpha\mu\nu\beta} = 0$$

Given these equations, R has 20 independent components (in 4 dimensions)

Flat manifold



$$R^{\alpha}_{\beta\mu\nu} = 0$$

Commutation of covariant derivatives and curvature

$$\nabla_{\beta} A^{\alpha} = A^{\alpha}_{;\beta} = A^{\alpha}_{,\beta} + \Gamma^{\alpha}_{\mu\beta} A^{\mu}$$

Now, let us do

$$\nabla_{\alpha} \nabla_{\beta} V^{\mu} = \nabla_{\alpha} V^{\mu}_{;\beta} = \left(V^{\mu}_{;\beta} \right)_{,\alpha} + V^{\sigma}_{;\beta} \Gamma^{\mu}_{\sigma\alpha} - V^{\mu}_{;\sigma} \Gamma^{\sigma}_{\beta\alpha}$$

In a locally inertial frame P, only the first term contribute, since the others will never contain derivatives of Γ :

$$\nabla_{\alpha} \nabla_{\beta} V^{\mu} (P) = \left(V^{\mu}_{;\beta} \right)_{,\alpha} = V^{\mu}_{,\beta\alpha} + \Gamma^{\mu}_{\nu\beta,\alpha} V^{\nu}$$

And also

$$\nabla_{\beta} \nabla_{\alpha} V^{\mu} (P) = V^{\mu}_{,\alpha\beta} + \Gamma^{\mu}_{\nu\alpha,\beta} V^{\nu}$$

The commutator

$$\left[\nabla_{\alpha}, \nabla_{\beta} \right] V^{\mu} (P) = V^{\nu} \left(\Gamma^{\mu}_{\nu\beta,\alpha} - \Gamma^{\mu}_{\nu\alpha,\beta} \right)$$

In P the Riemann tensor

$$R^{\alpha}_{\beta\mu\nu} = \Gamma^{\alpha}_{\beta\nu,\mu} - \Gamma^{\alpha}_{\beta\mu,\nu} + \Gamma^{\alpha}_{\sigma\mu} \Gamma^{\sigma}_{\beta\nu} - \Gamma^{\alpha}_{\sigma\nu} \Gamma^{\sigma}_{\beta\mu} \rightarrow_P \Gamma^{\alpha}_{\beta\nu,\mu} - \Gamma^{\alpha}_{\beta\mu,\nu}$$

so that, in the locally inertial frame P :

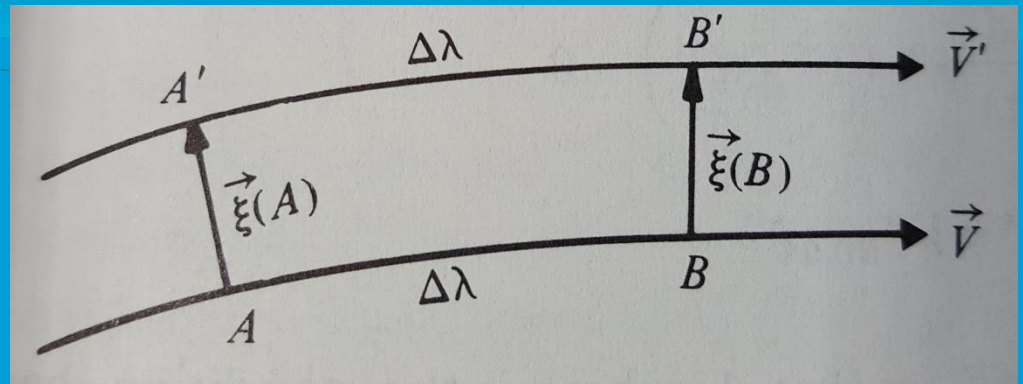
$$\left[\nabla_{\alpha}, \nabla_{\beta} \right] V^{\mu} (P) = R^{\mu}_{\nu\alpha\beta} V^{\nu}$$

This is a valid tensor equation \rightarrow valid in any coordinate system.

Geodesic Deviation

Parallel lines, when extended, do not remain parallel.

Now, suppose they start parallel at A and A':



Equation of the geodesics at A (assumed to be locally inertial), assuming that x^0 points along the geodesics ($V^\alpha = \delta^\alpha_0$)

$$\frac{d}{d\lambda} \frac{dx^\alpha}{d\lambda} \Big|_{(A)} = 0$$

Equation of the other geodesics at A', under the same assumptions (except local inertiality):

$$\frac{d}{d\lambda} \frac{dx^\alpha}{d\lambda} \Big|_{(A')} + \Gamma_{00}^\alpha(A') = 0$$

$$\Gamma_{00}^\alpha(A') \cong \Gamma_{00,\beta}^\alpha \xi^\beta$$

$$\frac{d}{d\lambda} \frac{dx^\alpha}{d\lambda} \Big|_{(A')} + \Gamma_{00,\beta}^\alpha \xi^\beta = 0$$

$$\frac{d^2 \xi^\alpha}{d\lambda^2} = \frac{d^2 x^\alpha}{d\lambda^2} \Big|_{(A')} - \frac{d^2 x^\alpha}{d\lambda^2} \Big|_{(A)} = -\Gamma_{00,\beta}^\alpha \xi^\beta$$

Now it is necessary to build the full second covariant derivative

$$\nabla_v \nabla_v \xi^\alpha = \frac{d}{d\lambda} (\nabla_v \xi^\alpha) + \Gamma_{\beta 0}^\alpha (\nabla_v \xi^\beta)$$

Now using the fact that in A: $\Gamma=0$

$$\nabla_v \nabla_v \xi^\alpha = \frac{d}{d\lambda} \left(\frac{d}{d\lambda} \xi^\alpha + \Gamma_{\beta 0}^\alpha \xi^\beta \right) + 0 = \frac{d^2}{d\lambda^2} \xi^\alpha + \Gamma$$

Bianchi Identity,
Ricci Tensor,
Einstein Tensor

Starting from the Riemann Tensor:

$$R^{\lambda}_{\beta\mu\nu} = \frac{1}{2} g^{\lambda\sigma} (g_{\sigma\nu,\beta\mu} - g_{\sigma\mu,\beta\nu} + g_{\beta\mu,\sigma\nu} - g_{\beta\nu,\sigma\mu})$$

$$R_{\alpha\beta\mu\nu} = g_{\alpha\lambda} R^{\lambda}_{\beta\mu\nu} = \frac{1}{2} g_{\alpha\lambda} g^{\lambda\sigma} (g_{\sigma\nu,\beta\mu} - g_{\sigma\mu,\beta\nu} + g_{\beta\mu,\sigma\nu} - g_{\beta\nu,\sigma\mu}) = \frac{1}{2} (g_{\alpha\nu,\beta\mu} - g_{\alpha\mu,\beta\nu} + g_{\beta\mu,\alpha\nu} - g_{\beta\nu,\alpha\mu})$$

And derivating:

$$R_{\alpha\beta\mu\nu,\lambda} = \frac{1}{2} (g_{\alpha\nu,\beta\mu\lambda} - g_{\alpha\mu,\beta\nu\lambda} + g_{\beta\mu,\alpha\nu\lambda} - g_{\beta\nu,\alpha\mu\lambda})$$

Using the symmetry properties of g and the commutation of partial derivatives

$$R_{\alpha\beta\mu\nu,\lambda} + R_{\alpha\beta\lambda\mu,\nu} + R_{\alpha\beta\nu\lambda,\mu} = 0$$

A valid tensor equation which is valid everywhere

$$R_{\alpha\beta\mu\nu;\lambda} + R_{\alpha\beta\lambda\mu;\nu} + R_{\alpha\beta\nu\lambda;\mu} = 0$$

Bianchi Identity

Ricci Tensor

$$R_{\alpha\beta} = R^{\mu}_{\alpha\mu\beta} = R_{\beta\alpha}$$

It is the only independent contraction of the Riemann Curvature Tensor

Ricci Scalar

$$R = g^{\mu\nu} R_{\mu\nu} = g^{\mu\nu} R^{\alpha}_{\mu\alpha\nu} = g^{\mu\nu} g^{\alpha\beta} R_{\alpha\mu\beta\nu}$$

Einstein Tensor

Starting with the Bianchi Identity

$$R_{\alpha\beta\mu\nu;\lambda} + R_{\alpha\beta\lambda\mu;\nu} + R_{\alpha\beta\nu\lambda;\mu} = 0$$

$$g^{\alpha\mu} \left(R_{\alpha\beta\mu\nu;\lambda} + R_{\alpha\beta\lambda\mu;\nu} + R_{\alpha\beta\nu\lambda;\mu} \right) = 0$$

this contraction yields

$$R^{\mu}_{\beta\mu\nu;\lambda} + R^{\mu}_{\beta\lambda\mu;\nu} + R^{\mu}_{\beta\nu\lambda;\mu} = 0$$

$$R_{\beta\nu;\lambda} - R_{\beta\lambda;\nu} + R^{\mu}_{\beta\nu\lambda;\mu} = 0$$

yet another contraction!

$$g^{\beta\nu} \left(R_{\beta\nu;\lambda} - R_{\beta\lambda;\nu} + R^{\mu}_{\beta\nu\lambda;\mu} \right) = R_{;\lambda} - R^{\nu}_{\lambda;\nu} + R^{\mu\nu}_{\nu\lambda;\mu} = R_{;\lambda} - R^{\mu}_{\lambda;\mu} - R^{\mu}_{\lambda;\mu} = 0$$

$$2R^{\mu}_{\lambda;\mu} - \delta^{\mu}_{\lambda} R_{;\mu} = 0$$

$$R^{\mu}_{\lambda;\mu} - \frac{1}{2} \delta^{\mu}_{\lambda} R_{;\mu} = 0$$

$$g^{\alpha\lambda} \left(R^{\mu}_{\lambda;\mu} - \frac{1}{2} \delta^{\mu}_{\lambda} R_{;\mu} \right) = 0$$

$$\left(R^{\mu\alpha} - \frac{1}{2} g^{\alpha\mu} R \right)_{;\mu} = 0$$

$$\left(R^{\alpha\mu} - \frac{1}{2} g^{\alpha\mu} R \right)_{;\mu} = 0$$

$$G^{\alpha\beta} = \left(R^{\alpha\beta} - \frac{1}{2} g^{\alpha\beta} R \right)$$

$$G_{;\beta}^{\alpha\beta} = 0$$

Einstein Equations for Weak Fields

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

$$|h_{\mu\nu}| \ll 1$$

Flat spacetime

curvature

Everywhere in spacetime

A coordinate system in which this happens is called a Nearly Lorentz Coordinate System

If one such system exists, then there are many others, linked by

- 1) Background Lorentz : $x^{\bar{\alpha}} = \Lambda^{\bar{\alpha}}_{\beta} x^{\beta}$ where the matrix Λ is a Special Relativity transformation. Under this transformation, one has $g_{\bar{\mu}\bar{\nu}} = \eta_{\bar{\mu}\bar{\nu}} + h_{\bar{\mu}\bar{\nu}}$ with h transforming as a tensor. This leads to the convenient fiction of thinking of a flat spacetime with a tensor defined on it.
- 2) Gauge transformations : small change of coordinate generated by a vector $x^{\alpha'} = x^{\alpha} + \xi^{\alpha}(x^{\beta})$ so that $g_{\alpha'\beta'} = \eta_{\alpha\beta} + h_{\alpha\beta} - \xi_{\alpha,\beta} - \xi_{\beta,\alpha}$ (the effect of coordinate changes is to redefine h).